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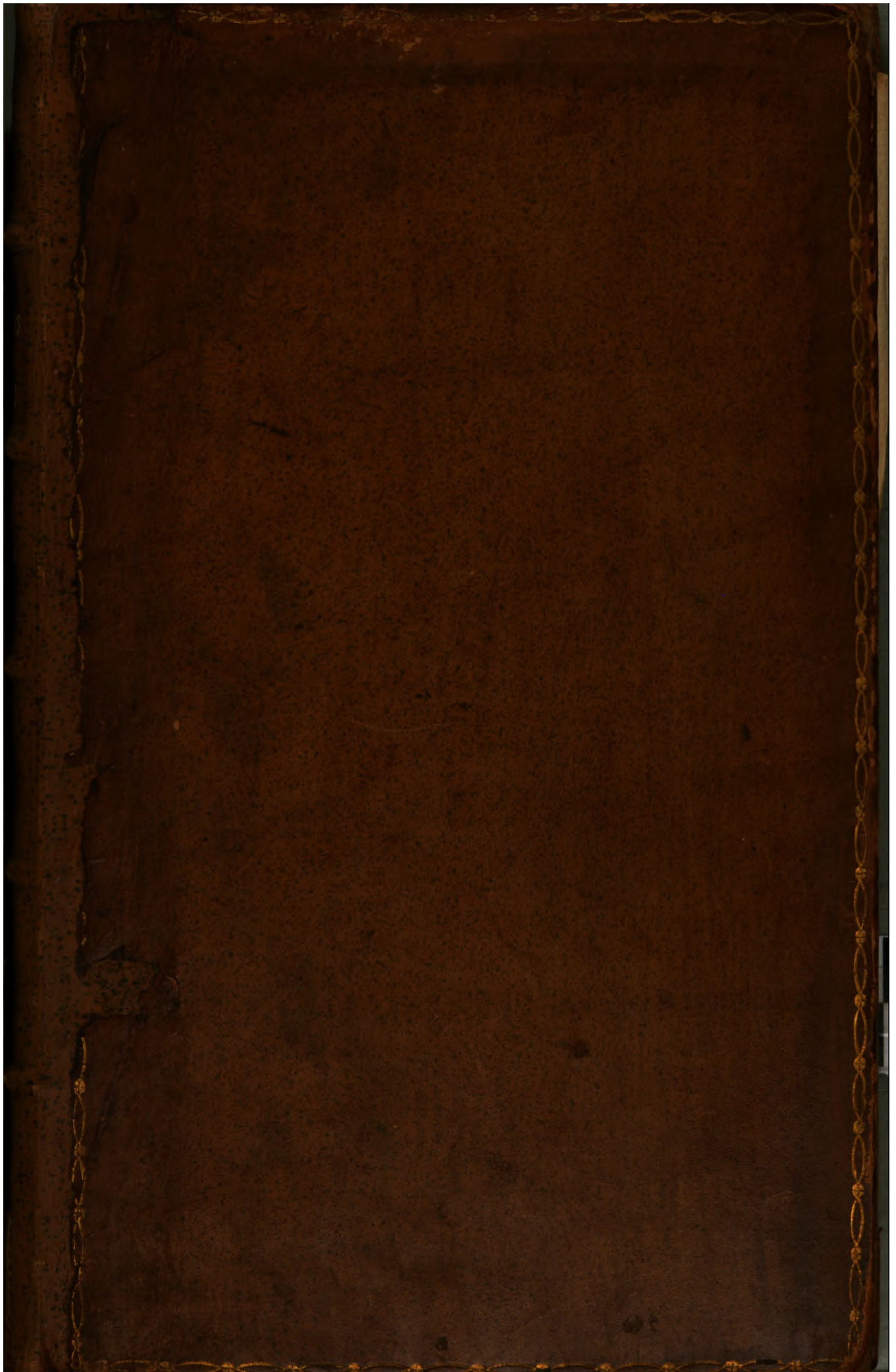
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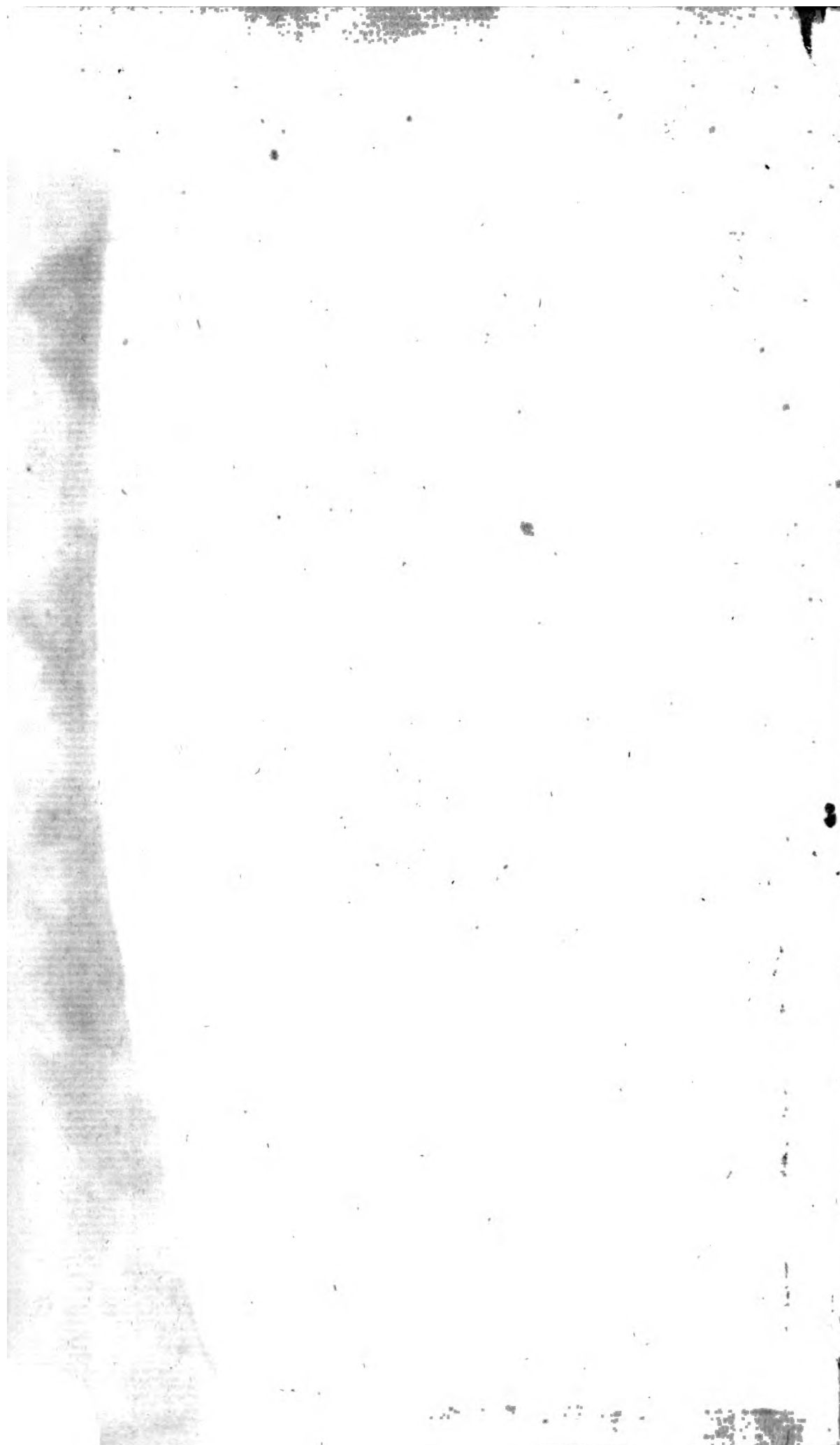


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S. P. Prigant
Sept. 30. 1834



Milnes

NEW
ELEMENTS of OPTICS;
OR, THE
THEORY of the ABERRATIONS,
DISSIPATION, and COLOURS
OF
L I G H T:
OF THE
General and Specific Refractive
POWERS and DENSITIES
OF
M E D I U M S;
THE
PROPERTIES of *Single and Compound*
L E N S E S:
AND
The NATURE, CONSTRUCTION, and Use
OF
Refracting and Reflecting
TELESCOPES and MICROSCOPES
Of every Sort hitherto published.

By *B. MARTIN.*

L O N D O N:

Printed for the AUTHOR, and sold at his Shop, the Sign
of the Globe and Visual-Glasses, two Doors below
Crane-Court, Fleet-Street.

MDCCLIX.

[Price Three Shillings sewed.]

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T H E

P R E F A C E.

IT is now about 20 Years since I publish'd a *Compendious SYSTEM OF OPTICS*; the general Reception it has been favour'd with by the Public, is one Reason why I have presumed once more to offer them my farther Thoughts on this interesting Subject. The Doctrine of *Refractions*, the Basis of this Science, I had not then thoroughly consider'd; I took it for granted, that there was nothing more to be known of it than what *Kepler, Des Cartes, Hugenius, Halley*, and Sir *Isaac Newton* had delivered. I did not know from their Writings, that there was a two-fold Refraction, *viz. a general and a specific Refraction* of Rays; and tho' I was always conversant with prismatic Experiments, yet I never adverted to the different Dimensions of the *colour'd Image* of the Sun, made by Prisms of different refractive Powers, nor observed what Proportion the Angle which those Images subtended bore to the Angle of total Refraction, supposing it to be constantly the same in Prisms of every Sort of Glafs.

And this seems to have been the general Case of other Opticians, till about the Year 1740. They all acquiescing in what Sir *Isaac Newton* had said in his Optics—*That Light in passing thro' different refracting Mediums, as often as by contrary Refractions it is so corrected, that it emergeth in Lines parallel to those in which it was incident, will ever appear white. But if the emergent Rays be inclined to the Incident, they will become*

tinged with Colours.—But these Positions are not universally true; for unless the Ratio of the Difference of Refraction in the most and least refrangible Rays to that of the whole Refraction of the Beam were constantly the same, there will be some Cases where *the emergent Rays may be parallel to the Incident, and yet be colour'd*; and on the other Hand, *the emergent Rays may be inclined to the incident ones, and yet may be white or colourless.*

And because this is extremely obvious in Experiments with the Prisms, it is amazing how it should escape the Notice of Sir *Isaac Newton*, who for near 70 Years made them his Amusement, and all other Opticians for many Years after; and still the more so, as it might easily be deduced from the Reason of Things.—For since whatever Degree of Refraction and Dissipation of Rays are produced by a given Incidence of Rays on a Prism of a given refracting Angle and Medium, the same can be corrected by another equal Prism of the same Kind, inverted; and that if the equal Prism be not of the same Medium, the said Refraction and Dissipation cannot be corrected, but there will be a Difference in both; and this Difference will be equal to the Refraction and Dissipation proper to a certain Incidence of Rays, and refracting Angle of the same heterogeneous Prism; and consequently if this latter be added to the former, it will correct the residual Refraction and Dissipation; and the Rays will emerge from the last Prism white or with Colours, but at the same Time will make an Angle with the incident Rays.

And from thence it will follow, that if Lenses analogous to such Prisms were properly consociated, they would have the same Effect; and so a compound Lens of a Telescope may be contrived that shall refract parallel Rays to the same Part of the Axis, or without Dissipation and Colours, as in common refracting Telescopes.

Nor

The P R E F A C E. v

Nor are we obliged to use Prisms or Lenses of a different Medium for this Purpose; Lenses of the same Sort of Glass will do, if we can by any Artifice contrive to have a different Degree of Refraction than what is natural in a given Incidence of Rays and Radius of Curvature in a Lens proposed. And tho' this Method be less obvious than the former, yet it was the first, and only one that has yet been made public.

All Writers on Optics have the common Complaint of refracting Telescopes, but are at the same Time all silent concerning any Method of remedying the same, at least any that have been genuine and effectual. Dr. HOOK contrived to *lengthen the focal Distance* of an Object-Glass, by combining a Lens of Water with one of Glass. Sir ISAAC NEWTON also proposes to correct the Errors of Refraction from the Figure of a Lens, by making a compound Lens with two Glasses and Water between them. But Dr. Gregory is the first I find mentioning any Method of correcting Errors from the different Refrangibility of heterogeneous Rays, by *an Object Lens of a different Medium*; but this he does in such a Manner, as plainly shews he had not a right Understanding of the Nature and Use of the Method he proposed, for he recommended this in Imitation of Nature, who, in the FABRIC of the EYE, has (says he) placed the *crystalline Humour* between the *aqueous* and *vitreous Humours*, (which are in a small Degree less refractive) in order that the Image may be painted as distinct as possible on the Bottom of the Eye. — But as the crystalline Humour is *convex*, and not *concave*, it cannot prevent the Dissipation of Rays in the Humours of the Eye; which Nature, by other Means, has most effectually provided against.

Other Expedients have been proposed by other Authors, but I have seen none that were properly adapted to answer the Purpose, till that of Mr. CALEB SMITH appeared,

pear'd, which was presented to the Royal Society, and publish'd in the 456th Number of the Transactions for January 1740, under the Title of, *A new Method of improving and perfecting catadioptrical Telescopes, by forming the Speculums of Glass instead of Metal.*

Mr. SMITH (among many other Things) assures us, this was his own Invention, in the following Words: — *Now, for this principal Defect (viz. the Errors from the different Refrangibility of Rays) no one, that we know of, has proposed any Remedy. — And as the Removal of so great an Impediment, is of the greatest Importance to the Science of DIOPTRICS, we have thought it worthy of a careful Examination, whether it might not be possible, in some Cases, for contrary Refractions so to correct each others Inequalities as to make their Difference regular; and if this could be conveniently effected, Sir ISAAC NEWTON has acknowledged there would be no farther Difficulty*.* — *Now upon due Consideration of this Subject, we have found it possible, by proper Methods and Expedients, to rectify those Errors which proceed from the different Refrangibility of Rays. — And our present Design is to shew what Advantage this will yield towards improving and perfecting catadioptric Telescopes, &c.*

The Author's Reasoning is founded on the well known and establish'd Principle, — *That the Sines of Refraction of Rays differently refrangible are to one another in a given Proportion, when their Sines of Incidence are equal.* Whence it is plain, this new Invention requires nothing more than the common Principles of Optics, properly applied, for its Explication.

He then proceeds to explain the ARTIFICE of his Invention, which, as to Theory, is equally ingenious and perfect. It consists in this; he shews how a *meniscus*

Lens

* *Philos Trans.* No. 88.

Lens may be used as a *Speculum*; and that the second Refraction in the LENS, which encreases the Error of the first, is, by the Mediation of a reflecting Surface, made to correct the said Error in the *Speculum* in a very great Degree, and the residual Error he corrects entirely by a concave Lens properly placed in the Axis; but as the Author has not given us the Procefs of his Demonstration, I have here supplied it, for the Satisfaction of inquisitive Readers.

As by this Contrivance the Refractions are made at concave Surfaces in the *Speculum*, it will moreover have this Advantage, that the same concave Lens which corrects the Error of *different Rays*, may at the same Time be made to correct the Error of the *Figure of the Glasses*, so far, at least, as to render it altogether insensible.

Therefore this Method is in itself quite perfect, and the only one that is so. For the other Method, which corrects the Errors of Rays differently refrangible by a convex and concave Lens, is extremely imperfect, as it makes the Errors from the Figure vastly greater than it is by the single Object Lens. And it is no Wonder if Mr. *Smith*, who consider'd this Affair principally in Theory, should give the Preference to the *catadioptric* Method; for he expressly mentions *other Methods*, by which the Errors of Refraction in common Telescopes may be remedied, and there is Reason to think he means the *different refractive Powers* of Glafs. In short, the two Methods are essentially the same, only the Mode or Way of applying the same Principle in Practice, is more perfect in one than in the other.

Now as the Impediment of refracting Telescopes has so greatly alarm'd the World of late Years, and as it proceeds from two Causes which are in general but little known; and lastly, as no Methods have been proposed, but this of Mr. *Smith*, for remedying the same, I have
thought

thought it would be acceptable to the Public to have the whole Doctrin explain'd from its first Principles, and applied in each particular Method as far as the Nature of the Thing required. I have endeavour'd to exhibit the Theory of each Sort of Aberration in the most natural and perspicuous Manner; and hope the Means of rectifying or remedying the same will be found in so clear and fair a Light, that every Reader will be able to form a Judgment of them according to their Merit; and probably many will conclude with myself, *That one Method is too difficult for Practice; and the other too imperfect to deserve much Regard from the judicious Part of the Public.*

Besides this I have given the Construction of all Kinds of TELESCOPES and MICROSCOPES, that have been made to the present Time, and illustrated the same by a great Number of Diagrams, to be found in no other Treatise of Optics; by which means Gentlemen may at once understand the Difference of Construction, as well as the *Rationale* and Merit of each particular Form. They will also be convinced, that there is Room enough for Improvement in refracting and reflecting Telescopes in many Respects, besides what relates to the *Aberration of Rays.*

To conclude, as I intend this for a Supplement to my SYSTEM OF OPTICS, in order to render the Theory of this Science compleat, I have not troubled the Reader with the Demonstration of the original Theorems over again, as they are there contain'd at large; but I have here given them a new Dress, and adapted them to express universally all the modern Improvements relative to all the Problems of optical Science. But the Demonstrations of the Theorems relating to Aberration of Rays, occasion'd by the Figure of *Specula* and *Lenses*, are to be found in my *Philosophia Britannica*; where the Reader will also find those above mentioned. Many Years have

The P R E F A C E. ix

have laps'd since this Treatise was first began, in all which Time I have seen nothing on these Subjects; and as I was obliged to proceed in a trackless Field, and many and difficult Experiments necessarily occur'd, my Progress was but slow; and it has been with much more Labour than any one would imagine, that I have at length finished this small Treatise, by which the Reader will be fully certified, that all the former and latter Improvements in Optics are owing (not to *Strangers* and *Foreigners*) but to *Englishmen* only, whose peculiar Glory it is to have an acknowledged Superiority, in regard to INVENTION, *over all the World besides.*



CONTENTS.

PART I.

- CHAP. I. *OF the different Sorts of RAYS of LIGHT, and Forms of LENSES used in TELESCOPES, with the Method of finding their focal Distances for PARALLEL RAYS.* Page I
- CHAP. II. *Of the focal Distances of converging and diverging Rays. Of forming the IMAGE of an Object in the Focus of a Lens, and on the Retina in the EYE; of the Position, Magnitude, and Distance of the Image; and of the visible Area, or Field of View.* 5
- CHAP. III. *Of the Structure of the Astronomical TELESCOPE, its magnifying Power and Use.* II
- CHAP. IV. *Of the Galilean TELESCOPE, and Opera-Glass; their Construction and Properties explained.* 15
- CHAP. V. *The Construction of the common Four-Glass TELESCOPE, and its Uses explained.* 19
- CHAP. VI. *The CONSTRUCTION and Rationale of the REFRACTING MICROSCOPE and MEGALASCOPE, Single and Compound.* 22
-

PART II.

- CHAP. I. *THE THEORY of the ABERRATION of RAYS caused by the FIGURE of Glasses, and the Indistinctness of VISION arising from thence, explained.* 26
- CHAP. II.

C O N T E N T S.

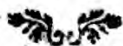
- CHAP. II. *The Method of correcting the Errors of this Aberration by a proper COMBINATION of Convex GLASSES; with a general THEORY for that Purpose.* Page 32
- CHAP. III. *The foregoing THEORY continued for a Computation of the Degree of Distinctness and Indistinctness of Vision in a proposed Combination of Glasses.* 36
- CHAP. IV. *The best DISPOSITION of Glasses in dioptric Instruments, according to the foregoing THEORY; and the different FORMS of TELESCOPES deduced from thence.* 44
- CHAP. V. *Of the Improvement and new Construction of compound MICROSCOPES, by the same Theory.* 50
-

P A R T III.

- CHAP. I. *THE general Doctrine of REFRACTION of Light explain'd; also the different Refrangibility and colorific Quality of heterogenceal RAYS.* 53
- CHAP. II. *The THEORY of the ABERRATION or DISSIPATION of heterogeneous LIGHT, explain'd and illustrated.* 60
- CHAP. III. *Of the general and specific REFRACTION, and DENSITY of Mediums; with the ANGLES of DISSIPATION, and THEOREMS for computing the same.* 63
- CHAP. IV. *The DENSITIES, and general and specific REFRACTIVE POWERS of different Sorts of GLASS, CRYSTAL, &c. determined by EXPERIMENTS.* 67
- CHAP. V. *Of the prismatic REFRACTION of LIGHT. THEOREMS determining the Angles of total REFRACTION and DISSIPATION of RAYS; with the METHODS of*

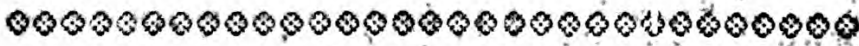
C O N T E N T S.

- of rendering the refracted Beam Colourless, tho' inclined to the incident one in any given Angle. Page 70*
- CHAP. VI. *Of the DOCTRINE of REFRACTIONS by LENSES of different Mediums, and the CONSTRUCTION of TELESCOPES with compound LENSES for preventing the DISSIPATION of Rays producing colour'd Light. 74*
- CHAP. VII. *THEOREMS for finding the focal Distances of compound LENSES, for any determined Lengths of TELESCOPES, OPERA'S, &c. 79*
- CHAP. VIII. *Remarks on the Nature, Construction, and Effects of the foregoing Achromatic TELESCOPE. 84*
- CHAP. IX. *The NATURE, CONSTRUCTION, and IMPROVEMENT of the CATADIOPTRIC or common REFLECTING TELESCOPE explain'd. 88*
- CHAP. X. *The Imperfection of VISION by the Achromatic TELESCOPE shewn by many EXPERIMENTS, and the great SUPERIORITY of the REFLECTING TELESCOPE evinced. 93*
- CHAP. XI. *The common THEOREMS of OPTICS tabulated, and rectified to express universally the focal Distance of RAYS reflected from MIRRORS, or refracted thro' LENSES of any Medium. 99*
- CHAP. XII. *An Illustration of the METHOD invented by Mr. CALEB SMITH, for correcting the ERRORS arising from the different Refrangibility of RAYS, by Means of a catadioptric SPECULUM and concave LENS. 107*
- CHAP. XIII. *LENSES made according to the Figures of the CONIC SECTIONS, or any other Curves, are subject to the same ERRORS from REFRACTION as those of a Spherical Form. 115*





T H E
 C O N S T R U C T I O N and U S E
 O F R E F R A C T I N G
 T E L E S C O P E S and M I C R O S C O P E S.




P A R T I.

Containing an Explanation of the Nature of VISION by LENSES, and the Structure and Properties of all the usual Forms of TELESCOPES; viz. the *Astronomical, Galilean, and Common Four-Glass Telescope*; and also of the Refracting MICROSCOPE.

C H A P. I.

Of the different Sorts of RAYS of LIGHT, and Forms of LENSES used in TELESCOPES, with the Method of finding their Focal Distances for PARALLEL RAYS.

1.  THE Word TELESCOPE is of Greek Original, and signifies the *Perfection of Vision*; and it is very justly so denominated, as it renders our View of distant Objects much more perfect than they appear to the naked Eye, and gives us a more certain Idea of what they are.

2. This noble Effect is producible two Ways. (1.) By *Refraction of Light* through *Lenses*; and (2.) by *Reflection of Rays of Light* from the polish'd Surfaces of *Mirrors*.

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The CONSTRUCTION and Use

3. The Rays of Light are of three Forms, *viz.*
 (1.) *Parallel Rays*, which proceed equally distant from each other; as those of the Sun, represented by ABCD. (2.) *Converging Rays*, which all tend in their Progress towards one Point; as the Rays of the Sun collected by the Glass BC, are made to converge towards the Point F. (3.) *Diverging Rays*, which proceed from any Point F, diverging or going wider of each other towards the Parts EG.

Fig. I.

4. The Point F, where the Rays of the Sun are collected into a small round Spot, is called the *Focus*, that is, the *Burning Point*; because the Rays being there condensed, the Heat is augmented, and they burn proportionably.

5. A **LENS** is any transparent Substance of a regular Form, through which Light may be transmitted, or refracted; those in common Use are of Glass ground and polished to a proper Figure, on which Account there are divers Sorts; as (1.) a *Plain Lens*, both whose Surfaces are parallel to each other. (2.) A *Plano-Convex*, consisting of one *Plain*, and the other a *Convex*, or *Spherical Surface*. (3.) A *Double-Convex*, both whose Surfaces are of a *Convex*, or spherically round Form. (4.) A *Plano-Concave*, with one *Plain*, the other a spherically hollow or *Concave Surface*. (5.) A *Double-Concave*, both whose Surfaces are spherically *Concave*. (6.) A *Meniscus*, which has one Surface *Convex*, the other *Concave*. Of which, and their Properties, in Order.

Fig. II.

6. RAYS OF LIGHT passing through a Plain Lens BC suffer no Alteration in respect of their Direction; for parallel Rays AB having passed through
 through

Of REFRACTING TELESCOPES.

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through the Lens, proceed on parallel towards D: And the same is to be said of the other Sorts of Rays.

7. If *Parallel Rays* AB, GH, ED, fall on the Convex Side of a Plano-Convex Lens BD, they will be converged to a Focus F, at nearly the Distance of the Diameter HF of the Sphere, of which the Lens is a Segment; or at the Distance of *twice the Radius* HC of the Sphere. Fig. III.

8. If Parallel Rays fall on a Double and equally Convex Lens BD, they will be conven'd to a Focus F, which will be in the Center C of the Sphere of Convexity, or distant in the first Surface H by the Radius HC of the Sphere, very nearly. Fig. IV.

9. If the Lens be not equally convex, parallel Rays will be converged to a Focus, whose Distance is easily assigned by the common Theory of Optics.

10. If *Parallel Rays* fall on a Plano-Concave Lens BD, they will be made to diverge afterwards towards IKL, as if they came from a Point F, nearly at the Distance of a Diameter of the Sphere of Concavity FH; which Point F is called the *Virtual Focus* of the Concave Lens. Fig. V.

11. *Parallel Rays* are made to diverge much more by passing through a Double Concave Lens; which, if it be equally concave on both Sides, will cause them so to diverge, as if they came nearly from the Center C, of the Sphere of Concavity. Fig. VI.

12. A MENISCUS LENS participates of the Properties, as well as the Forms of both the Convex and Concave Lenses. For if C be the Center of the Convexity BHD; then (1.) if the other Surface BID had been of the same Convexity the

B 2

Focus

The CONSTRUCTION and USE

4
Focus F of parallel Rays would have been at the said Center C, by *Art.* 8. (2.) If the said Surface had been plain, the Focus would have been at L, by *Art.* 7. (3.) But since, in the present Case, it is concave, the Focus will be removed farther to a Point F. (4.) If the said Surface BID were parallel to the other BHD, that is, if the Convexity and Concavity were equal, the Focus would be at an infinite Distance, or the Rays would proceed parallel after Refraction. (5.) If the Radius of Concavity were less than that of Convexity, the Rays would be made to diverge after passing thro' the Glass as in Concave Lenses.

13. From this Account of the Nature and Effect of Lenses, it appears, that if we look on F as a Radiant Point, or that from whence Rays of Light do proceed, then those Rays, if collected by any of the aforesaid Convex Lenses, placed at their respective focal Distances FH, will proceed parallel after they have passed through the said Lenses.

14. Also if Concave Lenses be made to receive converging Rays I, K, L, tending towards a Point F, which is at the Distance of the *Virtual Focus* of those Glasses from them, then those Rays, after having passed through the Lenses, will likewise proceed parallel to each other.

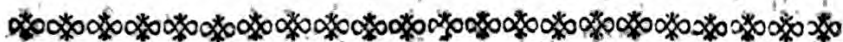
15. The most ready Way to find the *focal Distance* of any Convex or Meniscus Lens, is to hold it in the Sun-Beams, so that the Rays may fall perpendicularly upon it; then measure the Distance at which the Rays are collected in a white round Spot, well defined; for that is the Image of the Sun, Burning Point, or what we properly call the *Solar Focus* of the Lens.

16. To

Of REFRACTING TELESCOPES.

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16. To find the *Virtual Focus*, or Center of Concavity in any Concave Lens, proceed thus; cut a round Hole in a Piece of black Paper, &c. and lay it on the Concave Lens with a little Gum-Water, so that the Center of the Hole may be in the Middle of the Glass; then strike a Circle on another Piece of Past-Board, &c. whose Diameter is just twice the Length of that of the Hole on the Glass; then hold the Lens in the Sun Beams, and move the Paper backward and forwards, till the Rays diverge so much as just to fill the said largest Circle; then will the Distance between the Lens, and the said Paper be equal to the Distance of the *Virtual Focus* from the Lens.



C H A P. II.

Of the Focal Distances of Converging and Diverging Rays. Of forming the IMAGE of an Object in the Focus of a Lens, and on the Retina in the EYE; of the Position, Magnitude, and Distance of the Image; and of the Visible Area, or Field of View.

17. **S**INCE (as we have observed *Art. 7* and 8) *Parallel Rays* are converged to the Focus of a Concave Lens; it will be easy to conceive, that if the Incident Rays are not *Parallel*, but *Converging*, or *Diverging*, they will, in the former Case, be made to unite in a Point nearer

The CONSTRUCTION and USE

nearer to the Lens than the *Solar Focus*; and in the latter Case, they will be all conven'd in a Point beyond the said *Solar Focus* in the Axis of the Lens.

18. And because the *Radiant Point* must be at an *Infinite Distance*, in order that the incident Rays may be truly *parallel*; it must be at *more than an infinite Distance*, or be posited *negatively*, on the other Side of the Lens, that the incident Rays may be considered as *Converging*; and lastly, if the Distance of the Radiant be *less than infinite*; that is, if it be at any *finite Distance*, then the incident Rays will be *Diverging*, and have their *proper Focus* on the other Side the Lens, beyond the *Solar Focus*, as was said before.

19. And hence it follows, that the nearer the Distance of the Radiant is to Infinite, or the greater its Distance from the Lens, the less will be the Distance of the *Proper Focus* from the *Solar Focus*, and consequently from the Lens itself.

20. Thus as the *Radiant Point* approaches the Lens, the *Proper Focus* will recede from it; and, *vice versa*; so that if DEG be a Double and equally Convex Lens, the Centers of whose Convex Surfaces are *a* and *b* respectively; then if we consider ABC as an Object placed before it, and take therein the Point B in the Axis of the Lens, for the *Radiant Point*, or that from which we suppose Rays issue and fall diverging on the Lens; they will be collected all in the Point F on the other Side the Lens at a Distance EF, which is somewhat greater than the Distance of the *Solar Focus* Ea.

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Of REFRACTING TELESCOPES.

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21. In this Case, the Distance between the Solar Focus a , and the Proper Focus F is very small, because the Distance of the Object or Radiant Point B is remote from the Focus of the said Lens b .

22. But on the other hand, if we consider HFI as an Object placed near the Focus a of of the Lens, and F the Radiant Point; then because the Rays proceeding from thence to the Lens, must fall on it more diverging; therefore the proper Focus to which they will be converged on the other Side the Lens will be at B , at a great Distance from b the Solar Focus; because the Radiant Point is so near it.

23. If the Lens be a *Double and equally convex One*, and you know the Radius of Convexity, and Distance of the Radiant Point, you will find the Focal Distance of the Rays after Refraction through the Lens by the following easy Rule—*Multiply the Distance of the Radiant by the Radius of Convexity; divide the Product by the Difference between the said Distance and Radius; the Quotient will be the Focal Distance acquired.* Note, the Rule is the same for a Plano-Convex, if for *Radius* you take *twice the Radius*. (See *Art. 7 and 8.*) If the Lens be not equally convex, it will be difficult to find the Length of each Radius; and, when found, the Rule is too intricate for finding the proper Focus; which those who desire to know, may find in Page 96, of my *System of Optics*; or my *Philosophia Britannica*.

24. After the same Manner, if in the Object ABC we take either of the extreme Points A or C for the Radiant Point, then the Rays which proceed from thence to the Lens DEG , will by
it

it be converged to their respective Focus in the Points I or H, so that HFI shall be nearly in a Right Line, and equi-distant for the Lens DG.

25. It is further to be observed, that the Rays which proceed from the Points ABC in the Object, do in some wondrous and ineffable Manner convey or carry with them the Likeness or perfect Resemblance of those Parts of the Object to their focal Points HFI, where they appear as visible to the Eye as they do in the Object itself.

26. And since the same thing is to be understood of all the Points between A and C, it follows that there must be an adequate Representation of the Object ABC in the focal Points between H and I; and this we call the *Image* of the Object.

27. All those Rays which proceed from any single Point A, B, or C, in an Object, to the whole Surface of the Lens DG, is called (from its Figure) a *Cone*, or *Pencil of Rays*; of which, that which is in the Middle, is called the *Axis* of the Pencil, and passes through the Center E of the Lens, to the several Points in the Image IFH.

28. Hence it is evident, that the *Axes* of all the *Pencils of Rays*, issuing from every Point in the Object, meet and cross each other in the Center of the Lens; and consequently, that *Pencil of Rays* which comes from the Right-hand extreme Point C in the Object, will be all determined to the Point H in the Left-hand Extreme of the Image; and those which come from A will be made to represent the Point in
the

Of REFRACTING TELESCOPES.

9

the Image at I; from whence it is very evident, the *Position of the Image must be inverted, or contrary to that of the Object.*

29. 'Tis farther evident, that since the Axes of the Extreme Pencils of Rays AEI, and CEH, terminate the Length of the Object and Image, and cross each other in the Center of the Lens E, the Angles AEC and HEI will be equal; and so the Object AC, and its Image IH will be seen under equal Angles from the Center of the Lens.

30. For the same Reason also, we have the Length of the Object AC to the Length of the Image IH in the same Proportion as the Distance of the Object BE to the Distance of the Image EF. If the Object be a Surface, its Area will be to that of the Image, as the Square of BE to the Square of EF; and if it be a Solid, then the Solidity will be to that of the Image, as the Cube of those Lines BE and EF, respectively.

31. From what has been premised, it evidently follows, that if HI be the Object, then will AC be the Image formed thereof; so that the Object and Image are reciprocal, in a contrary Position, and proportional to the Distances from the Lens.

32. Hence also it is evident, that Objects AC, though large in themselves, if placed at a remote Distance from the Lens, will be represented in their Images IH very small, and very near the Focus a of the Lens; and therefore

33. If the focal Length Ea of the Lens, be increased or diminished, the Images of Objects
C will

The CONSTRUCTION and USE

will be proportionably increased or diminished. Or, in other Words, the *greater the focal Length of the Lens, the larger is the Image* formed thereby.

Fig. IX.

34. All the Pencils of Rays which proceed from an Object AC, placed in the Focus of a Convex Lens DG, do, after Refraction through it, proceed parallel (by *Art. 14.*) and if in this Case they are received by the Eye EFK, so as to pass through the Pupil PQ and the several Humours, they will by them be convened to their respective Focus's at the Bottom of the Eye, (*Art. 13.*) and there paint a lively Image IKH on the Retina; and thus *cause a distinct and perfect Vision of the Object ABC.*

35. If the two external Rays AD, CG, are conceived to proceed parallel to the Lens, they will be so refracted as to cross each other in the Focus of the Lens, which will be here near the Middle of the Pupil, at O, and contain an Angle DOG, which is called the *Optic Angle of the Lens*, and determine the Diameter of the *Visible circular Area, or Field of View*, as it is commonly called.

36. Since the Rays AD, CG, are supposed parallel, 'tis evident the visible Part AC of any Object, will be greater or lesser as the Diameter, or Aperture of the Lens, next the Eye, is so. Or, in other Words, the greater the Area of the Lens, the greater will be the *Visible Part, or Field of View*, in any Instrument that procures Vision by Means of a Convex Lens placed before the Eye.

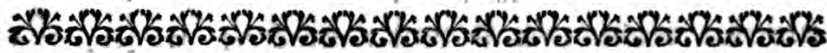
37. The

Of REFRACTING TELESCOPES.

11

37. The extreme Rays BD, BG, of any Pencil coming from any radiant Point B remotely situated, will, when they enter the Pupil of the Eye, or Lens DEG, be nearly co-incident with parallel Rays as AD, CG; and will therefore be fit for producing distinct Vision in the Eye, as having the Focus F nearly coincident with the Focus *a* of parallel Rays on the Retina at the Bottom of the Eye; as is plain to be seen in the Figure.

Fig. X.



C H A P. III.

Of the Structure of the Astronomical TELESCOPE, its Magnifying Power, and Use.

38. **W**E have premised such Principles, as are necessary to a right Understanding of the Nature, Construction, and Effects of *Refracting Telescopes*; which we shall next proceed to describe according to their several Kinds, beginning with those which are most simple and uncompounded; shew their several Properties, Powers of Magnifying, Methods of Improving them, &c.

39. The most simple Structure of a Telescope, is that which consists only of one Object Glass DEG, and one Eye Glass KLM; and fewer than these there cannot be for magnifying any distant Object ABC. For the Object Lens only serves to form the
C 2 Image

Fig. I.
Plate II.

The CONSTRUCTION and USE

Image HFI of the said Object, but does not magnify it, since the Eye placed at the Center E of the said Lens, views the Object and Image under equal Angles AEC and HEI (by *Art.* 29.) and consequently they must appear of equal Magnitude.

40. And since in this Telescope, it is the Image HI which we view, and not the Object AC itself, if we can contrive to view this Image under a larger Angle than that which the Object subtends, it will appear larger than the Object; and so the Object is in that Case said to be magnified. And this is easily effected, by fixing a Convex Lens KLM before the Eye, in such Manner, that the Image HI shall be exactly in the Focus of the said Lens. For then if we suppose the Rays IK , and IM to proceed parallel from the Image to the Lens KM , they will be convened to a Point O at the Pupil of the Eye, which shall be in the other Focus of the Lens (by *Art.* 7 and 8.) And since OL is equal to LF , and KM equal to HI , it is evident, the Angle KOM is equal to the Angle HLI , and is that under which the Image is viewed by the Eye through the Eye-Glass KM .

41. Now it is plain, the Angle HLI (or KOM) is to the Angle HEI , as EF is to FL ; because the Angle of apparent Magnitude of one and the same Object must always increase, as the Distance from the Eye decreases. Wherefore the Magnitude of the Image HI seen by an Eye at E , will be to its Magnitude seen from L as the Distance FL to the Distance FE . Or,

the

Of REFRACTING TELESCOPES.

13

the Power of Magnifying in this Telescope is in the Proportion of the Focal Length of the Object Glass FE to the Focal Length of the Eye-Glass FL.

42. For Example : Let the focal Length of the Object Glass EF be 5 Feet or 60 Inches, and that of the Eye-Glass LF equal 2 Inches ; then will the Magnitude of the Object seen by the naked Eye, be that of its Image seen by the Eye-Glass, as 2 to 60, that is, as 1 to 30; so that such a Telescope magnifies the Length of an Object 30 Times, or makes it appear as large as it would do, were it placed 30 Times nearer to the Eye.

43. And since the *Surfaces, or Areas* of Bodies, encrease in Proportion to the Squares of the *Lengths*, and the *Solidity* in Proportion to the Cubes thereof ; it follows that the *Superficies* of that Object is magnified (30 Times 30, or) 900 Times ; and the *Solidity* or *Bulk* 30 Times 30 Times 30, or) 27000 Times ; which is a very great Degree of Magnifying for an Instrument of 6 Feet Length. And this Power of Magnifying is encreased in Proportion as we encrease the focal Length of the Object Lens, as we shall shew more at large hereafter.

44. The *Object will appear inverted* in this Telescope (by *Art. 29.*) because it is the Image HI we view, which has a contrary Position to that of the Object AC. Also the *Visible Area, or Field of View*, is proportional to the Area of the Eye-Glass KM (*Art. 36, 37.*) And the Object (or, rather, the Image) will appear *more or less bright or luminous*, in Proportion as the Aperture, or Hole before the Object-Glass, is greater or lesser,

The CONSTRUCTION and Use

fer, because the Rays of Light, which enter thereby to form the Image, will encrease or decrease in that Proportion.

45. Since the Object, in this Structure of the Telescope, appears inverted, it is therefore unfit for viewing terrestrial Objects, which it turns topsy-turvy, and so makes the Sight of them unnatural, and consequently unpleasant. This Form or Structure is therefore wholly appropriated to astronomical Uses, in viewing the heavenly Bodies, whose Position we do not regard, since *Globes*, or *Circular Spaces*, appear to the same Advantage one Way as the other. In this Respect therefore it is called the **ASTRONOMICAL TELESCOPE**.

46. Since all the heavenly Bodies, except the Sun, are best viewed in the Night, the Use of the Tube (which in Part is to make every thing dark about the Image in the Day-time) is now unnecessary; and since most of the celestial Objects require a great Degree of Magnifying to render them very distinct and visible; therefore Object-Glasses of a very long focal Distance (above 100 Feet) have sometimes been used for this Purpose; which, as there was no Occasion for the Tube, were usually fixed on the Top of a very high Pole, &c. and an Eye-Glass properly adjusted thereto in the Axis; and both being thus fixed, they constitute what is commonly called the **AERIAL TELESCOPE**. But they are now out of Use since the same Powers of Magnifying are to be had in *Reflecting Telescopes* of short Lengths, and very easily applied to use.

C H A P. IV.

Of the Galilean TELESCOPE, and Opera-Glafs; Plate II.
 their Construction and Properties explained. Fig. II.

47. **I**N order to render the aforesaid Refracting *Telescope* fit for viewing terrestrial Objects, that is, for viewing them in an *erect Position*, a different Combination of Glasses has been introduced in two several Methods, *viz.* (1.) By using a *Concave Eye-Glass*; and (2.) by the *Addition of two Convex Eye-Glasses* to the former; so that a Telescope of this later Form, consists of an Object Glass, and three Convex Eye-Glasses.

48. The first Structure of these *Terrestrial Telescopes* is most simple; and, from its Inventor GALILEO, is called the *Galilean TELESCOPE*. The Effect of which Instrument is as follows: DEG is the Object-Glass, as in the other, but KLM is here a *Concave Eye-Glass* instead of the *Convex One*, used before. Now, all Rays AD, BE, CG, coming from an Object AC at a great Distance; and therefore parallel among themselves, or nearly so, will proceed (after Refraction thro' the Object Glass DG) converging towards its Focus, suppose at O, (by *Art. 7* and 8.) But if those Rays, instead of going to the said Focus O, should be intercepted by a concave Lens KM, placed at its focal Distance from the Point O, they will (after Refraction thro' it) pass *parallel*

rallel to each other, (by *Art. 15.*) and so are fit to produce distinct Vision (by *Art. 37.*)

49. Therefore the Object ABC will be *distinctly seen* by Means of the Concave Eye-Glass KM; and it will be seen *erect*, or in its *natural Position*; because the Rays, which come from the extreme Parts of the Object do not here cross each other, as before with the Convex Eye-Glass; by which Crossing of the Rays, the Inversion of the Object can only be effected, as has been already explained.

50. The Object will likewise appear *magnified*; for those Rays *aQG* and *bQD*, which cross each other in the Axis of the Object-Glass, at the Focus Q, will after passing the Lens, go parallel to the Concave Eye-Glass KM, by which they will be made to diverge so as if they came each from the Virtual Focus F, and will contain the Optic Angle KFM, under which the visible Part of the Object now appears to the Eye: But the Angle, under which the same visible Part (*viz. ab*) appears to the naked Eye, is the Angle *aQb*, which is equal to the Angle *DQG*; and since KM is equal to DG, (because the Rays DK and GM are parallel) therefore the Angle KFM will be as much greater than the Angle *DQG* as *QE* is greater than *FL*; and consequently, in the same *Proportion will the Object be as seen by this Telescope to its magnitude as seen by the naked Eye.*

51. Hence we observe, that if the Object Lens remain the same, and the Sphere of Concavity be equal to that of Convexity in the Eye-Glass,

Of REFRACTING TELESCOPES:

17

Glass, the magnifying Power will be the same, or equal in both (*Art. 42.*) And consequently, in this Case, the *Galilean Telescope* will be shorter than the other with a Convex Eye-Glass, by *double the focal Distance of the said Eye-Glass.* Or, on the other hand, if those two Telescopes are of equal Lengths, then *the Galilean will have the greatest Power of Magnifying*, as admitting an Object Glass whose focal Distance is longer than that of the other by double the focal Distance of the Eye-Glass, supposed the same in both.

52. Since only that Part of the Object will be visible which is terminated by those Rays which enter the Pupil of the Eye, 'tis evident, the *Visible Area*, or *Field of View*, in this Telescope, will be always proportional to the Area of the Pupil; and therefore must be much less than in the other Telescope, where the said *Visible Area* is proportional to the *Area of the Eye-Glass* (*Art. 36, 37, 45.*) And this is the great Reason why this is not the Telescope of most common Use.

53. For since it takes in so small a View, it proves very unsatisfactory, unless in short Lengths, which take in a greater Compass of Objects than larger Ones can do; for the shorter the focal Distance EQ , the greater will be the Optic Angle DQG to the naked Eye, if the Diameter of the Object Glass DG continue the same. Consequently the *shorter the Telescope*, the larger the *Field of View*, but the less will be the *Power of Magnifying*. And in this short Length they

D

are

The CONSTRUCTION and Use

are called OPERA GLASSES for their being used in the *Opera Houses*.

54. In regard to this Telescope, I shall only observe further, that the nearer the Eye is to the Lens KM, the greater Quantity of Rays it will take in, and consequently the greater will be the Field of View, and the more bright and luminous the Objects will appear; because the Rays of the several Pencils diverge, from the Eye-Glass, into a wider Space; and being denser near the Glass, will there enter the Eye in greatest Abundance, and produce the Effects above-mentioned.

55. In the last Place, as this Glass is very thin in the Middle, the Rays are less liable to be disturbed by the Irregularities in the Substance of the Glass, in passing through it; and also, since here we view not the Image of Objects, but the Objects themselves, by Rays coming directly from them, without crossing, or interfering with each other; it follows, that on both these Accounts, we have a clear and most distinct View of an Object by this Telescope. And hence it is, that we can sometimes see *Jupiter's Satellites* very plain, in a Telescope of this Sort, not above 20 Inches or two Feet long, when one of 4 or 5 Feet of the common Sort with Convex Eye-Glasses will scarcely render them visible.

C H A P.



C H A P. V.

The Construction of the Common Four-Glass TELESCOPE, and its Uses explained.

56. **W**E have shewn a Method of rectifying the Appearance of an Object by a Concave Eye-Glass, in the last Chapter; but as that is not accommodated for Use in the best Manner; we shall here treat of another, which makes the Structure of our common Refracting Telescopes, by the Addition of two Convex Eye-Glasses to that in the Telescope of *Art. 40.* and is as follows.

57. Since the Image HI is in the Focus of the first Eye-Glass KM, the Rays of every Pencil which come from every Point HFI therein, will, after Refraction through it, pass on parallel among themselves; and consequently, if they were intercepted by another Convex Lens OP, they would thereby be made to converge and unite in the Focus thereof, and there paint a second Image of the Object as QRS.

58. Now this *secondary Image* is to be formed in an *erect Position*, viz. the same with that of the Object; and, in order to that, the Distance of the two Lenses KM, and OP, ought to be equal to the Sum of the Focal Distance of both Glasses: For then the Axis of each Pencil of Rays HK and IM, which come from the extreme Points of the inverted Image HI, will

Plate. II.

Fig. III.

The CONSTRUCTION and USE

cross each other in N, the Focus of the Lens KM, and by that Means form an Image in a *contrary Position* in the Focus R of the Second Lens OP; which Image will therefore be *erect*.

59. As a Second Glass was necessary to rectify the Image, so a third Lens TV must be added to view the said Image thus rectified; and in order to have distinct Vision thereof, this third Glass must be placed at its focal Distance therefrom, because then the Rays coming from every Point in the Image QRS, will pass from the Glass parallel among themselves, and so enter the pupil of the Eye at W.

60. These parallel Rays are by Means of the several Humours of the Eye (but principally, the *Crystalline X*) convened to a Focus on the *Retina*, at the Bottom of the Eye, where a third Image YZ is formed of the Object ACB, and in an inverted Position; and thus distinct Vision of the Object is effected by Means of three Eye-Glasses. See *Art. 35.* and preceding.

61. And such is the Structure of the Refracting Telescope in common Use, in which all the three Eye-Glasses, KM, OP, TV, are of the same Convexity and Size, and therefore the focal Distance the same in all; whence the Distance between each, is double the said focal Distance; but it is not necessary that the Eye-Glasses should be of the same Convexity, or focal Length; for were they different in this Respect, yet if they were placed as above directed, *viz.* so that the Focus of each might fall upon the same Point N and R, the
Vision

Vision would still be distinct, and the magnifying Power of the Telescope only would be altered thereby.

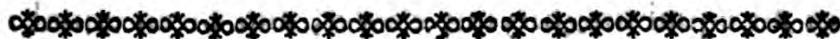
62. If all the Eye-Glasses be of the same Sphere or Convexity, as they commonly are, then will the *magnifying Power*, or *Field of View*, be the same as it would be with only one of them, KM, as is easy to conceive from a View of the Scheme. For the Images QRS and HFI are equal to each other, and viewed by Glasses of an equal magnifying Power, and of an equal Area or Size; and therefore the Appearance of the Object must be every way the same by three as by one of them, the *Position* only excepted.

63. And this is the most commodious Structure of a Refracting Telescope, where the Objects to be viewed are large and well enlightened, as the *Sun* and *Moon* among the heavenly Bodies; and the *People, Houses, Trees, and Fields*, (if not at too great a Distance) among terrestrial Objects. But very small Objects, as the Planets, and especially their Secondaries or Moons, require too nice and distinct a Vision to be effected by three Eye-Glasses, (in short Telescopes) since each Glass reflects a great deal of Light which is thereby lost, and the Refraction through three Glasses causes the Vision to be much less perfect and distinct than it would be through one only.

64. But two Things are here to be observed for the greater Perspicuity and Distinctness of Vision in this Sort of Telescopes, *viz.* (1.) That the Image QS or HI, which ever is seen, be duly circumscribed with a proper Diaphragm, that is with a Piece of Wood with a Hole in the
Middle,

The CONSTRUCTION and Use

Middle, to exclude or cut off all extraneous or useless Rays, such as are refracted from the Sides or Edges of the Lenses. (2.) That the Eye be placed at W, so that the focal Point of the Eye-Glass may fall as near as possible on the Pupil of the Eye; for else the Area or Field of View will not be so great, and what is still worse, the Object will appear coloured and confused towards the extreme Parts.



C H A P. VI.

The CONSTRUCTION and Rationale of the REFRACTING MICROSCOPE and MEGALASCOPE, Single and Compound.

Plate I.

Fig. XI.

65. **T**HE Nature and Construction of a *Refracting* MICROSCOPE being similar to that of a Telescope, will require but little to be said to explain them. They are indeed essentially the same Thing, and differ only in *Mode* and *Accidents*. The same Thing is also to be observed of that Instrument we properly call a MEGALASCOPE, as being of a middle Nature between both the former.

Fig. VIII.

66. We have already observed, that if ABC be regarded as a distinct Object, then IFH will be the Image thereof formed near the Focus *a* of the Glass *g e d*, which in that Case was considered as the Object-Glass of the Telescope, (*Art. 39.*) But on the other Hand, if IFH be

Of REFRACTING TELESCOPES.

23

a small Object, placed near to the Focus (*a*) of a Lens *gd*, whose focal Distance *ea* is very small; then will an Image AC be formed thereby of the Object IH, which will be larger than the Object in the Proportion of *eB* to *eF*; and in this Case the Glass *gd* is the Object-Lens of a MICROSCOPE.

67. But this Lens may be applied two Ways to magnify the Appearance of an Object, *viz.* (1.) By itself, and then it is called a *Single Microscope*; or, (2.) in Construction with another, and then it is called a *Compound Microscope*.

Plate I.

68. It is evident from what we have said, (*Art. 13.*) that an Object placed in the Focus *F* of any Lens *BD*, will be seen distinctly by the Eye placed near to it, because the Rays from every Point of that Object will be refracted parallel to the Eye, which is all that is necessary for distinct Vision (37.)

Fig. { III
IV

69. Now it is known by Experience, that no Object can be seen by the naked Eye at a less Distance than *six Inches*, generally; therefore as much as the focal Distance *FH* of a small Lens is less than six Inches, so many Times will the Object be magnified, that is, it will appear under an Angle so much larger than what it subtends to the naked Eye; since it is a well known Proposition in Perspective, *that the apparent lineal Dimensions of Objects are as the Optic Angles under which they are seen by the Eye.*

70. Hence an Object placed in the Focus of a Lens at the Distance of *one Inch*, will appear through it six Times longer and broader than

it

The CONSTRUCTION *and* USE

it can be viewed by the Eye, and therefore 36 Times larger in *Surface*, and 216 Times in *Bulk* or *Solidity*.

71. Again, suppose the focal Distance of a *Spherule*, or very small Lens, were only $\frac{1}{10}$ of an Inch; then is the Object view'd by it at a Distance 60 Times less than six Inches, and consequently will appear magnified 60 Times in *Length and Breadth*, 3600 Times in *Surface*, and 21600 Times in *Bulk*, or solid Content.

72. When the focal Distance is smaller, the magnifying Power is as much greater in Proportion, and some Spherules have been made, whose focal Distances have not exceeded the 40th or 50th Part of an Inch. Such Glasses discover exceedingly small Objects, and yet some of them appear but as animated Points, and require yet greater Powers of Magnifying to shew their Parts distinctly.

73. When the focal Distances of Glasses are less than *an Inch*, they are properly called MICROSCOPES, as they are fitted to render the *smallest Objects visible*.

74. But Glasses whose focal Lengths are from one Inch to six, may be esteem'd MEGALASCOPIES in different Degrees, as they magnify all the *larger Sort of small Objects*, and by this Means afford a much more beautiful and satisfactory View of such Objects than can be had by the Eye alone.

75. In the second Way of magnifying small Objects, we use two Glasses, *viz.* an Object-Glass *d e f*, which makes a large Image *ABC* of a small Object *a c b* placed near its Focus (*f*). This Image *AC* is viewed distinctly through the
Eye-

Plote I.

Fig. XI.

Eye-Glass DEF placed at its focal Distance EB, and under the Angle DIF by the Eye at I; whence it is evident, that this Compound Microscope in *Fig. XI*, is nothing more than the two *Fig. VIII* and *IX* put together. And that whereas in a *Single Microscope* we view the Object itself immediately; in the Compound one, we view the Object ab only in its magnified Image AC.

76. The Power of Magnifying in a Compound Microscope, is thus easily understood; let the Distance of the Object ec from the Lens df, be *one Inch*, and the Distance of the Image AC be eB six Inches; then if we view the Image AC with the bare Eye, we shall see the Object therein *six Times magnified*: But if we view this Image through the Lens DF placed at its focal Distance EB one Inch, then will itself appear six Times larger than before to the Eye alone, and consequently the Object will appear six Times six, or 36 Times magnified. Or the Angle DIF (equal to CEA) will be 36 Times bigger than that which the Object ab alone subtends to the naked Eye at the Distance of six Inches.



P A R T II.

Containing the THEORY of the *Aberration* of Rays arising from the Figure of the Glasses; the *Indistinctness* of VISION occasioned thereby; the Means of correcting the same by a proper *Combination of Lenses*; and of the NEW FORMS of TELESCOPES derived from thence, consisting of *three, four, five, and six Glasses*, with their *most advantageous Positions*.

C H A P. I.

The THEORY of the ABERRATION of RAYS caused by the FIGURE of Glasses, and the Indistinctness of VISION arising from thence, explained.

77. THE Telescopes hitherto described for viewing Objects erect, consist of four Glasses only; and we find no other in Use to the Times of Mr. *William Molyneux* here, and *Hugenius* abroad. But neither of these Authors have given us a Print or Diagram of a Telescope with any more than four Glasses, tho' each of them has considered the Nature of Vision by two Convex Lenses instead of one; and *Hugens* only has mentioned the Advantage of such a Combination. What *Molyneux* did was only to find the focal Distance of two such Glasses

Glasses which he received from Mr. *Flamsteed*, in a Letter dated Jan. 17, 1689.

78. The Use of two Glasses is to correct or Plate II.
diminish the Error of one single Glass arising from its Spherical Figure, therefore it will be first necessary to explain what this Error is, and then how it is in a great Measure corrected by two Glasses. The first who gave the Rationale thereof was Sir *Isaac Newton*, whose Method of explaining this Matter I have shewn at large in my *Philos. Britannica*, from the following Diagram, where NBM is the Spherical Surface Fig. IV.
of a Plano-Convex Lens NGMB; C the Center; CB the Radius or Semi-diameter taken in the Axis; AN an incident Ray, and NK the same refracted, cutting the Axis produced in the Point K. Also let F be the Focus of parallel Rays which pass through the Glass infinitely near to the Axis; let FD be a Perpendicular to the Axis in the Point F, then will KF be the Error, or Difference of the Focal Distance of parallel Rays which are incident near the Axis and at the Distance GN, the Semi-aperture of the Lens. This is called the *Aberration of the extreme Ray in Longitude*.

79. Again, let any Ray (*an*) be incident on the other Side the Lens, at the Distance *bG*, the refracted Part of this Ray *nd* will intersect the other refracted Ray ND in the Point Q, at the Perpendicular Distance QO from the Axis; this is called the *Lateral Error*, or the *Aberration in Latitude*.

80. It is evident from the Figure, that as the Ray (*an*) approaches the extreme Ray AM,

The THEORY of the ABERRATION of RAYS,
 the Point of Interfection Q will approach the Axis; and when an coincides with AM , the Point Q will coincide with the Point K in the Axis; and it is as obvious that Point Q will coincide with F, when the Ray (an) approaching the Axis aB , at last becomes coincident with it: Therefore there is *one Position* of the Ray (an) in which it will cut the Ray ND in a Point Q, which will make QO a *Maximum*, or the greatest of all.

81. If we take the Arch $Bm=Bn$, and $BM=BN$, the Rays incident on m and M will intersect in the Point P on the other Side, and so make $PQ=2QO$; and it is also plain (1.) That all the Rays which fall on the Lens between N and M are refracted through the Space PQ. (2.) That PQ is the Diameter of the least Circular Space possible, in which all the Rays can be congregated, because there will be some Ray (an) that will meet the extreme Ray ND, at the Distance $QO=\frac{1}{2}QP$ from the Axis. (3.) Hence it follows, that this circular Space is the *Focus or Place of the Image* of an Object belonging to parallel Rays incident on the Lens NM.

82. Further, by reason of similar Triangles KOQ, KFD, and NGK, we have $QO : KO :: DF : KF :: NG : GK$. But it is demonstrated, † that when QO is greatest, then $KO=\frac{1}{4}KF$, and also that KF is always $\frac{2}{3}$ of GB, the Thickness of the Lens; so then $KO=\frac{2}{3}GB$, and consequently $GK : GN :: \frac{2}{3}GB : QO$, whence $\frac{9GB \times GN}{8GK} = OQ$; whence PQ, the Diameter

of

† Philof. Britannica.

of the Circle of Aberration, is known for any given Lens.

83. It is also demonstrated, † that the Error PQ will always be proportioned to $\frac{NG^3}{BC^2}$; so that when the Radius is given, *the Error will be as the Cube of the Aperture directly: And when the Aperture is given, the said Error will be as the Square of the Radius inversely.*

84. It is also to be demonstrated, that when the Convex Side of the Lens NBM is turned towards parallel Rays, the Error KF will be but $\frac{7}{8}$ of the Thickness of the Lens GB, and *therefore near four Times less than in the other Case; for $\frac{7}{8}GB : \frac{3}{8}GB :: 54 : 14$, which is almost as 4 to 1.*

85. If the Side NGM, instead of being a Plane, were a Spherical Surface described, with a Radius equal to six Times the Radius CB, then the Error KF would still be less; for then it would be $\frac{15}{14}$ of the Thickness of the Lens. But if this Lens were turned with the least convex Side to the distant Radiant, the Error would be $\frac{145}{42}$ of the Thickness. In the first Position it is the best Glass, and in the last, the worst of all others; yet in the best Position, it differs so little from the *Plano-Convex*, as not to deserve notice.

86. If the Side NGM were Concave, so that the Lens became a *Meniscus*, there is no Proportion of the Radii, or Position of the Lens with regard to the Radiant, but what will give the Aberration KF greater than in the *Plano-Convex* in its best Position. And since this was first observed by Opticians, the *Meniscus* began to
loose

† Ibid.

lose Ground in the Construction of Optical Instruments, and is now quite rejected, tho' formerly in such high Estimation.

87. Lastly, it is demonstrated, that the Aberration PQ is as the Square of the Sine of Refraction (the Sine of Incidence being Unity) in all Mediums of different refractive Powers: Thus if a Lens of the same focal Distance and Aperture were made of *Glass and Water*, and suppose those Sines in *Glass* to be as R to 1, and in *Water* as r to 1; then will PQ in the *Glass* Lens be to the same in the *Water* Lens as R^2 to r^2 , or the Area of the Circles of Aberration, and of Course the Indistinctness of the Object will be as the Refractions R and r of the Mediums.

88. Whatever has been observ'd with regard to *Convex* and *Plano-Convex Lenses*, will hold good in *Concave* and *Plano-Concave ones*. And in both Sorts, it is supposed that all of them have the same focal Distances, Apertures, and Thicknesses, while we are comparing their respective Aberrations.

Fig. V.

89. After the same Manner Sir *Isaac Newton* has shewn how to estimate the Aberration arising from the Figure of the Speculum NBM, by reflected Light. Thus let CB = Radius, FB = $\frac{1}{2}$ CB = Distance of the principal or solar Focus F; and AN, *an*, two incident Rays parallel to the Axis, and ND, *nd*, the same reflected, intersecting each other in Q, and the Axis in K and *k*, as by refraction before. Also let NG be the Sine, and GB the versed Sine of the Angle of Incidence ANK or NKB.

90. Then

90. Then it is demonstrated, * *That the Longitudinal Aberration* $FK = \frac{1}{2}GB$ *nearly*; and therefore always *proportion'd to the versed Sine of Incidence* GB , or to the *Square of the Sine* NG , because GB is as NG^2 constantly.

91. Again, when QO is a Maximum, then $KO = \frac{1}{2}KF$; and because KG is nearly equal to BF , therefore $GK:GN :: KO:QO = \frac{GN \times KO}{GK}$; whence the Value of PQ (the lateral Error) is known, because $KO = \frac{1}{2}KF = \frac{1}{2}GB$.

92. Lastly, it is shewn that this lateral Error PQ is always *as the Cube of* NG *the Semi-aperture of the Speculum*, when the Radius is given; and *as the Square of Radius inversely*, when the Aperture of the Mirror is given: That is, universally, PQ is as $\frac{NG^3}{CB^2}$ (as at *Art.* 83.)

93. From what has been said, it is very evident, that if Rays proceed from any Point, as (*a*) at an infinite Distance to a Lens NM , the Image of that Point will not be a Point, but the *Area of a Circle*, whose Diameter is PQ ; and therefore that Point (*a*) cannot be distinctly represented; but will be rendered *indistinct and confused in Proportion to the Area of the said Circle of Aberration* in the Lens or Speculum, as it is the Image of this Circle (or *dilated Point*) that is impressed on the *Retina*, and excites the Idea of the Point (*a*) in the Mind.

Fig. VI.

94. Hence it appears, that the Points in the Surfaces or Substances of Bodies cannot be perfectly and distinctly seen, as each of them will be dilated into a sensible Area; and such as are

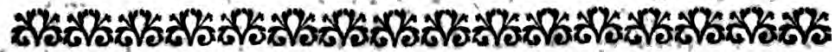
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* *Philos. Britannica.*

The ERRORS of ABERRATION corrected

contiguous (as *a, b, c,*) will have their confus'd Images all blended together nearly in the same Space, *viz.* the *Circle of Aberration*, whose Diameter is *PQ*.

95. Hence the Stars, which, as to Sense, are only lucid Points; will appear to have some Magnitude (and not as Points) in the Focus of the best Sort of Telescopes, even supposing there were no other Cause of Confusion or indistinct Vision besides what resulted from the *Spherical Figure of the Lens or Speculum*.



C H A P. II.

The Method of correcting the ERRORS of this Aberration by a proper COMBINATION of Convex GLASSES; with a general THEORY for that Purpose.

96. **A**S the indistinct Vision of Objects is so sensible and considerable in single Lenses, and increases in proportion as the *Square of the Radius* decreases, (83) it is no Wonder if the Learned in Opticks have employ'd their Skill to correct it as much as possible, in order to render Dioptric Instruments more perfect. Several Ways have been essay'd for this Purpose, but none so successfully as that by using a Composition of *two Lenses* instead of *one*; and the first I find describing this Method is the celebrated *Hugenius* in his *Dioptrica Nova*, who has said but very little on the Subject; nor do I find

at all consider'd by any late Writer in Optics; and, to the best of my Remembrance, I have seen no Diagrams of Telescopes where this Contrivance has been applied, but that one of the above-named Author, and what has been copied from it.

97. But as much depends on this Doctrine, I shall here deliver the general Theory, and then apply it in the Construction of Telescopes: Suppose parallel Rays AN, BM, fall on a Lens MN, and are refracted to a Focus at F; if at any Distance less than CF another Convex-Lens GH be placed to receive the said refracted Rays, it will converge them sooner to some Point *f*, which is call'd the *Compound Focus* of both the Lenses. The Optic Angle G *f* H is equal to that form'd by some single Lens EF, whose focal Distance is Q *f*.

Fig. VII.

98. Let $\left\{ \begin{array}{l} CF = F \\ OP = y \\ Qf = x \end{array} \right\}$ be the focal Distance of the Lens $\left\{ \begin{array}{l} NM \\ GH \\ EF \end{array} \right\}$
 $\left\{ \begin{array}{l} Of = f, \text{ the Compound Focal Distance.} \\ CO = D, \text{ the Distance of the Lenses} \\ \text{NM and GH.} \end{array} \right.$

99. The Rays from the first Lens NM fall *converging* on the second GH, tending to the Point F; if now we put $OF = d$, we shall have by common Optics $d = \frac{yf}{y-f} = F - D$; from which Equation we get $F + y - D : F - D :: y : f$. And thus the *compound focal Distance* *Of* is known.

100. After the same Manner the parallel Rays LG, SH are refracted by the Lens GH
F to

The ERRORS of ABERRATION corrected

to the Lens NM tending to a Point y ; but are converged by it to a Point f , which is the *compound Focus* on that Side; and now it is

$$Cy = \frac{Ff}{F-f} = y-D; \text{ whence } F+y-D : y-D ::$$

$F : f = Cf$, which therefore is known also.

101. Because of the similar Triangles FNC, FGO, and fEQ , fGO ; and $EQ = NC$; we have $CF : OF :: NC (=EQ) : GO :: Qf : Of$; that is, $F : F-D :: x : f$. Whence

$$\frac{Ff}{F-D} = x.$$

102. But we had above (99.) $f = \frac{F-D \times y}{F+y-D}$;

which substituted for f in the last Article will

give $x = \frac{Fy}{F+y-D}$; from which Theorem we

have a Solution of the following *Problem*, viz.

103. *Given the focal Lengths of two Lenses, and the Distance between them; to find the focal Length of a single Lens that has the same magnifying Power, or Optic Angle, with the two combined ones.*

By this RULE,

Divide the *Product* of the two focal Lengths of the given Lenses, by their *Sum* less'n'd by their *Distance*, and the *Quotient* will be the focal Length of the single Lens as required.

104. From that general Theorem, the Analyst will easily deduce the following Solutions, viz.

105. Given F , x , and D , to find y ? The

Theorem is $y = \frac{Fx - xD}{F - x}$.

106. Also,

by a COMBINATION of LENSES:

106. Also, Given y , x , and D , to find F ?

The Theorem is $F = \frac{x D - x y}{x - y}$.

107. Lastly, Given F , y , and x , to find D ?

The Rule is $D = F + y - \frac{F y}{x}$.

108. From these Theorems, and others to be deduced from them, by supposing $D = 0$, $F = y$, $F = D = y$; or supposing F , y , or f , to be negative, &c. any thing relating to a Comparison between a Combination of two Lenses and a single one in respect of the magnifying Power, Distinctness of Vision, &c. may easily be determined, whether for Eye-Glasses or Object-Glasses, or both, as we shall see further on.

109. We may also observe from the foregoing Analogies (99, 101.) that we have the Ratio of the *compound focal Distances* $O f$ and $C f$, thus $f : f :: \overline{F - D} \times y : y - D \times F$; and therefore when $f = f$, then $F = y$, or the said focal Distances can never be equal but when the Lenses are so.

110. Lastly, we may observe, that since the parallel Rays LG SH , refracted thro' both the combined Lenses, intersect the Axis in the same Point f , as it would do if it were refracted by the single Lens EF , as is evident by continuing it to R ; therefore since $GO = RQ$ it will follow, that the Diameter IK of the principal Pencil of Rays KfI diverging from the Focus f will be the same as it would be if it proceeded directly to the single Lens EF ; and consequently this Combination of Lenses makes no Alteration in that Respect.

Scholium.

111. From what has been said of a Combination of Lenses, it appears that if we consider GH as a *Constant Lens*, and NM as a *variable one*, then (1.) if F be the Focus of *parallel Rays* falling on GH, those Rays will all be collected into that Point (7.) (2.) The Addition of the Lens NM, if *Convex*, will cause the *parallel Rays* AN, BM, to become *converging* ones NG, MH, which therefore by the Lens GH will be refracted to a Focus f , so that the focal Distance $O f$ will be less than OF. (3.) If NM be *Concave*, the *parallel Rays* will be refracted thro' it diverging on GH, by which Means they will be refracted beyond the Point F to some Focus P; all which is evident from common Optics. (4.) In each Case the *Aber-ration from the principal Focus will be as the Square of OF to the Square of O f or OP respectively*; as is plainly demonstrated by *Hugenius* in his Diop. Pag. 247. but would be too tedious here to transcribe.



C H A P. III.

The foregoing THEORY continued for a Computation of the Degree of Distinctness and Indistinctness of Vision in a proposed Combination of Glasses.

112. **W**E have seen the Effects of a Combination of two Glasses in regard to the *magnifying Power*, we shall now examine according to the Theory what the Case will be with respect

respect to the *Distinctness of Vision*, that is to say, what the *Distinctness of the View* will be thro' *two Lenses* compared with that thro' *one* of the *same magnifying Power*.

113. In order to this it must be consider'd, that since the *Aberration, or Indistinctness of Vision*, is as the *Square of Radius inversely*, (83.) the *Distinctness of Vision* will be as the *Square of Radius directly*; and therefore if by Means of *two Glasses we can get the View of an Object*, where the *Radii of the Glasses* bear a greater Proportion to their respective Apertures, than the *Radius of a single Glass of equal magnifying Power* does to its Aperture, it is evident the *Distinctness of that View* will be promoted in Proportion to the Square of that Ratio.

114. For Example: Suppose $F-D=y$, or OF to be the focal Distance of the Lens GH , so that the Focus of each of the Lenses NM and GH falls on the same Point F ; then by the Theorem (102) we have $x=\frac{1}{2}F$, or $Qf=\frac{1}{2}CF$: Also, since in this Case we have $F:y::x:f$ (101) therefore $f=\frac{1}{2}y$, or $Of=\frac{1}{2}OF$. Now since we have the *same Optic Angle GfO* by both the Glasses, as by the single one EF , the Ratio of the Radius OF to the Aperture GO , or of the Radius CF to the Aperture NC , is double the Ratio of Of to OG , or of the Radius Qf to the Aperture EQ , and therefore the *Distinctness of Vision* by both the Lenses is *four Times greater than that by the single Lens EF*.

115. The same Thing may be demonstrated from the Consideration that the Aberration PQ
is

The THEORY of Distinct VISION,

is in the same Glass always proportion'd to the Cube of the Semi-aperture EQ , or Sine of half the Optic Angle EfQ , (see Art. 83.) and that in small Angles (as in the Object Glasses of Telescopes, &c.) the Sine EQ is nearly as the Angle EfQ . The Aberration therefore being as the Cube of that Angle, it is plain if we make the same Angle by two Refractions instead of one, the Quantity of the Aberration will be greatly lessen'd, since the Sum of the Cubes of the Parts will be much less than the Cube of the whole; and when the Parts are equal, the Sum of Cubes of each will be but a fourth Part of the Cube of the whole. Thus if the whole Angle EfQ be as 1, the Cube thereof is 1; but the Half is $\frac{1}{2}$, the Cube of which is $\frac{1}{8}$, and twice that $\frac{2}{8} = \frac{1}{4}$, which is as the Aberration arising from the two Halves, and is therefore but a fourth Part of the whole.

Fig. VII.

116. This is evidently the Case when the Optic Angle $GfO (=EfQ)$ is made by two Refractions, by the two Lenses NM and GH , so posited that the Focus of each may fall on the same Point F ; for then the Angle $GfO = LGf$, which is composed of the two Angles LGF , and FGf , the first of which $LGF = TNF = GFO$, (by reason of the parallel Lines TN , LG , and FC) which is the Part made by the Lens NM . Also the Angle FGf is the Refraction of the Ray NG , or second Refraction of the Ray AN ; and since in the present Case, $Of = fF$ (as we have shewn 114.) and Of in small Angles is equal to Gf nearly; therefore the Angle GFO is equal to the Angle FGf

FGf very nearly, those Angles being in the same Ratio with the equal Lines Gf and fF , when they are not large; and the Optic Angle $GfO = GFO + FGf$; consequently the *Aberration* PQ is but $\frac{1}{4}$ Part so great by the two Lenses NM and GH together, as it is by Lens EF alone.

117. But to render this Theory general for any Position or Form of the Lenses NM and GH , it is evident, since the Aberration is lessen'd by dividing the Optic Angle, the Distinctness of Vision will be thereby promoted; and because each of the Angles contribute thereto in Proportion to its Magnitude, the joint Effect of both Parts, or Angles GFf and fGF , will be as the Product or Rectangle under both, or as the Rectangle of the Lines Og and fF ; but according to our former Notation (98.) $Ff = F - D - f$; and $Og = f$; consequently $Ff - Df - ff$ will be every where as the *Distinctness of Vision* by the two Lenses above that of a single Lens, of the same magnifying Power.

118. Let the Degree of Distinctness thus obtained be represented by $G = Ff - Df - ff$; when this is a *Maximum*, or the greatest possible, the Fluxions thereof will be nothing, viz. $Ff - Df - 2ff = 0$, whence $F - D = 2f$; or $FO = 2Of$, that is $Og = fF$, or the Angle $GFf = fGF$, in the best Position of the Lenses, as before demonstrated.

119. Consequently since in that Position we have shewn the Distinctness of Vision is 4 Times as great as by a single Lens (116.) that will be the whole Effect of a Combination of two Glasses, and it may be shewn that 3 Glasses will produce 9 Times the Distinctness; and so on in Proportion

The THEORY of Distinct VISION,

tion to the Square of the Number of Glasses; but then if we consider the Evil to be remedied is but small, and the Damage we sustain in Loss of Light and Irregularity of Refraction thro' so many Lenses, we may soon make the Remedy worse than the Disease; and every thing consider'd, it will appear *that two Lenses are better than a greater Number.*

120. We shall now apply this Theory in some practical Cases, and illustrate them by Example. In order to this, let the focal Lengths of the Lens NM be $F=3$, that of the Lens GH be $y=2$, and then in their best Position we have also $F-D=y=2$; whence the focal Length of the equal Lens GH will be (102.) $x=$

$$\frac{Fy}{F \times y - D} = \frac{Fy}{2y} = \frac{1}{2}F = 1,5 = Qf.$$

121. In the Example above, the Distance of the Lenses $CO = F - y = D = 1$; if this Distance be increased, both the magnifying Power and Distinctness will be diminished; thus suppose $D=2$, then $x=2$: But as the focal Length of the Lens EF encreases, the magnifying Power of the Telescope decreases. Also $f = \frac{F \times x - D \times x}{F}$ (99.) $= \frac{2}{3}$; therefore $\overline{F-D-f} \times f = G = \frac{2}{9}$, which is the Degree of Distinctness; and is less than $\frac{2}{9}$ or $\frac{1}{4}$ (when $D=1$) in the Proportion of 8 to 9.

122. Again, suppose $D=3$, then $F-D=0$ and $x=F$; but in this Case the compound focal Distance $f = \frac{F-D \times x}{F} = 0$; consequently the Distinctness $\overline{F-D-f} \times f = G = 0$; so that the Advantage gain'd

gain'd by the Combination of two Lenses here, ceases in that Respect also.

123. On the other Hand, if both the Lenses are placed together, so that $D=0$; then the focal Distance of the *equal Lens* will be $x = \frac{Fy}{F+y} = 1,2$; the *compound focal Distance* will be $f = x = 2$: Also the Distinctness will be as $F-f \times f = G$ (117.) but $F-f : f :: \frac{3}{5} : \frac{2}{5}$ (taking $F-D=F=1$) therefore $F-f \times f = \frac{3}{5} \times \frac{2}{5} = \frac{6}{25} = G$, the Distinctness in this Position, which is not much less than in the best, where $D=1$, and the Distinctness as $\frac{1}{4} = \frac{6}{24}$. Wherefore the Ratio in this Case and that, is as 24 to 25. But 24 is nearer to 25 than 8 is to 9, (as in (121.) where $D=2$) and consequently this Position is better than that.

124. Again it is evident, from the Theorem $x = \frac{Fy}{F+y-D}$, that however you vary the Magnitude of y , (provided it does not exceed F) neither the magnifying Power or Distinctness of Vision will be alter'd thereby in the best Position of those Glasses, because since then $F-D=y$, it will ever be $x = \frac{1}{2}F$, a constant Quantity; and because in this Case y or OF will ever be bisected by the compound Focus f , the Distinctness will ever be the same, and a *Maximum*, (119.)

125. In what has been said, Regard has been had to the Vision of an Object RS placed in the compound Focus f exterior to the Glasses; but it makes no odds, if that Focus be supposed *negative*, or the Image *between* the Glasses, which will always be the Case when D is greater than

The THEORY of Distinct VISION,

y . Thus let $F=3$, $D=2$, $y=1$, then $x=1,5$ as before, and the Degree of Distinctness also the same, since $F-D=y$.

126. If the focal Distance of one Lens be supposed infinite, as F , then the Lens NM becomes a plain Glass, and the Composition ceases, and consequently the Distinctness of Vision occasioned thereby now vanishes also.

127. If one of the Lenses be concave as NM , then will its focal Distance be negative, or $-F$; and the Theorem (102.) becomes $x = \frac{-F y}{y - F - D}$, for the Focus of the equal Lens EF , which will always be *positive* while $F + D$ exceeds y ; but otherwise *negative*, unless $y = F + D$, then it becomes *infinite*, or the Rays are refracted parallel to each other from the Lens GH .

128. The compound focal Distance is in this Case $f = \frac{-F x - D x}{-F}$, (99.) and consequently must always be *affirmative* while x is so. And when $D=0$, then $f=x$.

129. The Expression for the Distinctness of Vision (118.) now becomes wholly negative or $-F f - D f - f f = G$, or changing all the Signs, $F f + D f + f f = -G$; but *negative Distinctness is Indistinctness*, therefore *Objects are viewed to a Disadvantage by Convex and Concave Lenses combined together*.

130. For Example: Let the Concave NM have a negative focal Distance $-F=3$, also let $y=2$, and their Distance $D=1$; then will the focal Distance of the equal Lens EF be $x = \frac{-6}{2-4} = 3$, and the compound focal Distance

$f =$

$f = \frac{-12}{-3} = 4$. Whence in this Case $FO = 2fO$, which is just the Reverse of the Case where F is positive, or the Lens NM convex, (118.)

131. If $-F=y$, and $D=0$, that is, if a Concave and Convex Lens be placed together, and their focal Distance equal, then x becomes infinite, or the Rays after Refraction will be parallel as before.

132. But if two such Lenses are placed at a Distance from each other, they will have a magnifying Power equal to that of some Lens EF .

For Example: Let $-F=3$, $y=3$, and $D=1$; then the focal Distance $x = \frac{-9}{-1} = 9$; and the compound focal Length is $f = \frac{-18}{-3} = 6$.

133. A Concave and Convex Lens placed together of different focal Lengths, will have a magnifying Power. Thus, suppose $-F=3$, $y=2$, and $D=0$; then $x = \frac{-6}{-1} = 6$. And in this Case $f=x=6$ also.

134. From these Instances it appears, that a Convex and Concave Lens combined together, always magnify less than the Convex Lens alone, and ever diminishes the Distinctness, and therefore ought by no Means to be used but when some particular Circumstances render such a Combination necessary.

Scholium.

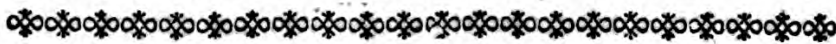
135. From what has been said it evidently appears, that the compound focal Distance of the combined Lenses is nothing more than the focal Dif-

The best DISPOSITION of GLASSES

Distance f found by the common Theorem of Optics $\frac{dr}{d-r}=f$, adapted to the constant Lens GH, where $OF=r$, and $OP=f$, when the Rays are *diverging*; or $Of = -f (= \frac{-dr}{-d-r})$ when the Rays fall converging on the Lens GH.

136. Since the Aberration is ever as OF^2 to Of^2 , or OP^2 ; (111.) therefore when $F=3$, and $y=2$, and $D=1$, we have OF^2 to Of^2 , as 4 to 1; so the Aberration in the Focus f is 4 Times less than in F , agreeable to (116.)

137. But if it be $-F=3$, $y=2$, and $D=0$; then $OF : OP :: 2 : 6$; and therefore $OF^2 : OP^2 :: 4 : 36 :: 1 : 9$. Whence the Aberration in the compound Focus f will be to that in the Focus P , as 1 to 36; or the Aberration will be 36 Times greater with the Concave, than with the Convex combined with the Lens GH, in the Cases now specified.



C H A P. IV.

The best DISPOSITION of Glasses in Dioptric Instruments, according to the foregoing THEORY; and the different FORMS of TELESCOPES deduced from thence.

138. **A**FTER having thus particularly stated and illustrated the Theory of distinct and indistinct Vision, so far as it depends on the Figure of the Glasses, and is to be remedied by
a Com-

a *Combination* or rather *Duplicature* of them; there will be no Occasion to spend much Time in the Application of it to the Improvement of our common refracting Telescopes.

139. For with regard to the Glasses themselves, it is evident this Theory may be applied as well to *Object-Glasses* as *Eye-Glasses*; for when we supposed $F=3$, and $y=2$, and $D=1$, &c. if these Numbers denote Inches, they are applicable to *Eye-Glasses*; but if they denote Feet, then they will have Relation to *Object-Glasses*, for in this Case $F=36$, $y=24$, and $D=12$ Inches. And therefore if the Telescope be constructed with two Object-Glasses, whose focal Lengths are 36 and 24 Inches, and placed at the Distance of 12 Inches, they will magnify as much as a single Glass whose focal Length is $x=15$ Inches, or Half that of the greatest alone, (124.)

140. Then with regard to the Position of the Glasses of either Sort, we have shewn that it is indifferent whether they are placed together or asunder, provided they are so placed that their Focus's coincide or fall on the same Point; for in all such Cases the magnifying Power will not be alter'd; and the Distinctness of Vision will be the greatest possible, (118.) viz. 4 Times as great as by the *single Lens* of the same *magnifying Power*.

141. Where there are but three Glasses in a Telescope, if the Image be between the two Glasses next the Eye, and their Distance D be equal to, or exceeds F the focal Distance of the largest, then the said largest Lens is to be consider'd as
com-

combined with the Object-Glass; in all other Cases it is combin'd with the Glass next the Eye, and this is a Circumstance necessary to be adverted to, as it makes a considerable Difference in the magnifying Power, as will appear by the following Example.

142. Let the focal Distance of the Object-Glass NM be = 15 Inches, that of the middle Glass GH = 3; the Distance $D = 12$; then will the Image be equal to that made by a single Lens of 7,5 Inches focal Distance: And if the Eye-Glass AB be 1,5 Inch, the Power of magnifying will be as 5 to 1. Whereas if the middle Glass be combined with the Glass near the Eye in the best Manner, the magnifying Power will then be as 15 to 1,5, or as 10 to 1, which is double of the former.

Plate III.

Fig. I.

143. But if the two Glasses next the Eye, viz. AB and GH, be so posited that their Focuses may coincide, and the Distance between them exceed the focal Distance of the lesser Glass AB, then will the Glass GH contribute to form the Image IK more perfectly than could be done by the Object-Glass NM alone; and also, to correct the Error in viewing it by a single Eye-Glass AB. Therefore this must be the best Disposition of Glasses, for an *Astronomical Telescope*.

Fig. II.

144. To form the Telescope for viewing *Terrestrial Objects* in the best Manner, that is, to shew them upright and as free from the Aberration of the Figure of the Glasses as possible, to the three Glasses NM, GH, and IQ of the *Astronomical Telescope*, (143.) we must add three others, viz. TR, OP, and AB; the first of these TR is to be

be of the same Size and focal Length with LQ , and placed at twice the Distance of the compound Focus S from the said Glafs; and what that Distance is, we have shewn how to find. By this Means a second Image VW will be formed of the first in a contrary Position to it, and therefore the same with that of the Object itself, as is evident from (58.)

145. This Image VW is view'd to the greatest Advantage also by two Glasses OP and AB , as we have so often shewn from the Theory. And by the two Glasses extraordinary GH and OP , the *Field of View* is considerably enlarged, and the Use of the Instrument made so much the more agreeable. For the Lens GH contracts the Image of each particular Object, and therefore brings many more into the Field, which are view'd more perfectly by the Glafs OP in conjunction with AB , than by the Lens AB alone. All which is evident from the same Principles.

146. But since this Disposition of Glasses is very different from that of the *Six-Glafs Telescope* commonly sold in the Shops, it will be necessary to shew the Reason why we place them in this Manner, and this it will be easy to comprehend from the following Diagram.

147. For let IM be the Image formed by the Object-Glafs, and IA a Ray incident on the first Eye-Glafs AB ; this Ray will be one of those belonging to a Pencil proceeding from the Point I , and whose Axis is the Ray IG passing thro' the Center F of the Lens AB .

Plate III,
Fig. III.

148. Now let the Point I be the Focus of Rays parallel to the Axis IG, and infinitely near to it; and let Ad be another Ray parallel to the said Axis, but incident on the Glass at the Distance AF, then by the Theory it appears (78.) that the Focus of such a Ray will be some Point (a) in the Axis nearer to the Glass than the Point I; so that Ia will be the Aberration of such a Ray.

149. Therefore some Ray DA inclined to the Axis IG and diverging from it, will fall on the Glass AB in the Point A, and be refracted to the Point I. Therefore, *vice versa*, the Ray IA will be refracted to its proper Focus in its own Axis IG produced; the Distance of which from the Glass AB may be easily found by the Theorem $\frac{d r}{d-r} = f$ in common Optics, because a I will always be equal to $\frac{1}{3}$ of the Thickness of the Lens AB, if it be equally convex, that is $Ia = \frac{1}{3} FR$. But FE the geometrical focal Distance of the Lens is known, and therefore $FI = d$, and consequently $Fa = r$, the focal Distance of parallel Rays dA ; whence $f =$ focal Distance belonging to the Ray IA after Refraction is known of Course.

150. The Ray AD in its Progress falls on the second Lens CD, which we suppose to be equally convex, or the same with the Lens AB, and placed at double the focal Distance ($HO = OR$) from it. Let QH be drawn thro' H (the Center of the Lens CD) parallel to IG, then shall this Line continued be the Axis proper to the refracted Ray DL, which it will intersect in L

151. Now because LQ is parallel to IG, and the Ray AD is equally inclined to both, it must at its Incidence at A and D suffer an equal Refraction in all Respects, and therefore DL will be equal and parallel to AI: Also HD is equal to AF, and PL equal to EI. Consequently the Point I in the Image IM is represented at L in the same Situation and Distance, with respect to the Lens CD and its Axis PH. And because what we said with regard to the Ray IM, and the Point I, is equally demonstrable of every other Point in the Image IM, it follows, that the second Image LK is *exactly similar*, and *every way equal to the first Image*, but in an *inverted Position*.

152. Hence it is plain, that the Error of Refraction from the Figure of the Glass AB is adequately corrected by the second Glass CD, and that the Image LK is, so far, a perfect Representation of IM. And it is not in the Power of Art by Glasses to make an Image *more perfect than the Object*, or to give a more distinct View of an Object in its Image, than we have of itself when view'd under the same Angle, even supposing there were nothing in the Glass to hinder the most perfect Refraction of Light.

153. If now we suppose the Glass CD not similar, and equally convex with the Glass AB, it will follow, that the Ray AD will not be alike and equally refracted by it; and that the Aberration occasion'd by the Glass AB cannot be adequately corrected by the Lens CD in a con-
H
trary

50 *Of the IMPROVEMENT and New CONSTRUCTION*
trary Refraction, and therefore that the second
Image LK cannot be equal to, nor equally
perfect with the first IM.



C H A P. V.

Of the Improvement and New Construction of
Compound MICROSCOPES, *by the same*
Theory.

154. **A**S the Aberration from the Figure of
the Glasses required a Correction in
the *refracting Telescope*, so does it also in the
compound Microscope, and in a much greater De-
gree, since the more Convex a Lens is, and the
greater the Distance between the geometrical and
and proper Focus, the greater will this Kind of
Error prove; and therefore in the Microscope,
where both these Points are in an extreme De-
gree, there will be required the utmost Assistance
from Art to correct it.

Plate III.

Fig. IV.

155. This appear'd so sensibly the Case, even
to former Opticians, that they were soon directed
to add to the single Eye-Glass AB another and
larger Lens EF, by which Means the Vision of
the Object OQ was render'd more perfect in the
Image IM, and the Field of View at the same
Time enlarged, as in the Telescope.

156. But

156. But notwithstanding this Improvement by the Addition of a third Lens to the Microscope is so very obvious both in Practice and Theory, yet *Hugénius*, who has consider'd this Sort of Aberration in Microscopes more particularly than any one, has not given any Iconism of this Instrument with more than two Glasses. The same may be said of Mr. *Molyneux*, Dr. *Gregory*, and most other Authors; who seem wholly unapprized of the Advantage of this Sort of Construction, in Point of *Distinctness of Vision*, at least they say nothing at all of the Matter.

157. It is not only evident from the Theory of this Aberration, that the Image of any Point is render'd less confused by Refraction thro' *two Lenses*, than by an equal Refraction thro' one; but also it follows from the same Principle, that the same Point has its Image still less confused when form'd by Rays refracted thro' *three Lenses*, than by an equal Refraction thro' two; and therefore a third Lens GH added to the former two, will contribute to make the Image more distinct, and consequently the Instrument more compleat.

158. At the same Time the Field of View is amplified; and the Use of the Microscope render'd more agreeable by the Addition of the Glass GH, and that very sensibly above what is to be found in the *common Construction of a Three-Glass Microscope*.

159. Since the *Distinctness of Vision* is thus promoted by the Glass GH, we may allow a somewhat larger Aperture to the Object Lens

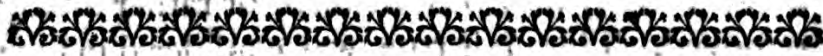
A Four-Glass MICROSCOPE.

KI; and thereby encrease the Brightness of Objects, and greatly heighten the Pleasure of viewing them.

160. And tho' each of these Particulars is easily confirm'd by Experiments, as well as demonstrable from Theory, yet are all the Optic Writers totally silent with respect to a *Four-Glass Microscope*. And as there is no mention of such a Thing, there is no *Iconism* of it of Course; and therefore this which I have here given, will prove at least a Novelty to the Reader in Print.

161. For I do not pretend to say no Microscope with four Glasses has been made, I know they have, and it is more than 30 Years since I bought one of the same Construction nearly with what I have here given; and it excell'd by much any Three-Glass Microscope I ever saw; but could not learn who the Maker was, it being an old Instrument when I first had it. It had its Origin therefore from Experience, and not from Theory; as in all Probability it was made before *Hugenius's* *Dioptrics* was publish'd, which is the first Book that contains the Theory of the Aberration from the Figure of Glasses that I have heard of.

162. As to the best Manner in which these four Glasses are to be placed in the Microscope, there can be no Difficulty to those who understand the Theory; and those who do not, it does not at all concern. But both in the Telescope and Microscope there will always be *once particular Disposition or Construction* of a given Number of Glasses that will be the best of all others, and it is the Excellency of a *Theory*, that by it alone such an interesting Point can be determined.



P A R T III.

Containing the THEORY of the *Aberration* of RAYS arising from their *different Refrangibility*; and the Method of removing the same by a proper COMBINATION of LENSES of *different Figures* and *refractive Powers*, demonstrated from the Nature of the *prismatic Refraction* of Light. With Theorems for finding the focal Distances of *double* or *triple* COMPOUND LENSES; and their Application to TELESCOPES and MICROSCOPES.

C H A P. I.

The general Doctrine of REFRACTION of Light explain'd; also the different Refrangibility and colorific Quality of heterogeneal RAYS.

163. **W**E are now to consider the *second Source* or Cause of indistinct or imperfect Vision in *dioptric Instruments*, viz. that which results from the *different Refrangibility* of the Rays of Light, in their Passage thro' *Media* of different refractive Powers; and of the proper Methods by which the *Aberration* of Rays, and the Imperfection of dioptric Vision occasion'd

tion'd thereby may be remedied, or rather totally removed. This, it is true, was consider'd in such a Light by the great Author of this Part of Philosophy, as by him to be thought impracticable; but Experience often convinces us that the wisest Men may be mistaken, and the Annihilation of this Sort of Aberration or Error is so far from being impossible or desperate; that it may be effected more Ways than one, as has been discover'd more than 20 Years since; and which we shall now proceed to explain, having first premised the following Principles, by way of *Data*, as they are confirm'd by the most certain and obvious Experiments.

164. *First*, That the Particles of all Matter are affected with certain Powers, by which they are *attracted and repelled*.

165. *Secondly*, That these Powers are of greater Energy in the Particles of some Sorts of Matter or Bodies than in others.

166. *Thirdly*, That this Power acts with the greatest Force upon the Surface of Bodies, and decreases not simply as the Distance increases, but according to some Power of the Distance.

167. *Fourthly*, That it is the same, or acts with the same Force, at equal Distances from the Surface.

168. *Fifthly*, That a Ray of Light is real Matter, or that the Particles of Light are only Particles of Matter attenuated or subtilized to an extreme Degree.

169. *Sixthly*, That the Particles of Light and all other Matter mutually act upon each other by Attraction and Repulsion.

170. *Seventhly*, That the Force of any Body is (*cæteris paribus*) proportion'd to the Number of Particles it contains, or to its *Density*.

171. *Eighthly*, That this Force is (*cæteris paribus*) porportionate in anyBody to the *intensive or peculiar Force of its Particles*.

172. *Ninthly*, That therefore the absolute Force of any Body is proportional to the *Density* thereof, and to the *specific Force* of its Particles conjointly.

173. *Tenthly*, That the same Particles of Matter in some (if not in all) Bodies are affected with *different Forces on their different or opposite Sides*. For Satisfaction in this Particular, see Sir *Isaac Newton's* Optical Queries relating to *Island Crystal*.

174. These Principles, as I said, are indisputably true by different and numerous Experiments; on these, therefore, we now proceed to explain the various Phænomena of refracted Light and Vision occasioned thereby; observing the Consequences, and making the necessary Deductions as we go along, for explaining what relates to the *Distinctness* and *Indistinctness of Vision* by *dioptric Instruments*, and particularly *Telescopes*, and how far they are capable of Improvement from this Doctrine.

175. Let ABC be a Section of the *Blade* of a *Knife*, held with its Edge a little below a Beam of the Sun's Light, *abc*; and it appears by Experiment, that the Rays *Aa* which pass nearest the Edge of the Knife, are inflected towards the Side thereof to (*d*); and those which are farther from the Edge, are less bent out of their proper Direc-

Fig. V.

Plate III.

Directions, till at last the attracting Force ceasing at the Distance of the Ray (be) that Ray passes strait on: But immediately beyond, the Rays are repell'd by the Knife, and from the Direction cg , they are inflected the other way towards f . So that by both these Powers the Beam of parallel Rays is made to spread and diverge after it has pass'd by the Edge of the Knife. See much more on this Head in Sir Isaac Newton's Optics, and Gravesande's Elements, &c.

Fig. VI.

176. Let there be two contiguous Mediums zt and sy separated by a plain Surface st ; and suppose sy the densest Medium of the two: Let the Distance, to which the attracting Power of its Particles extends, be denoted by sq ; and take $sv = sq$; and thro' q and v draw qr and vw parallel to st ; then is $qrwv$ called the Space of Attraction.

177. Now suppose a Ray of Light were incident from the Medium zt on the Medium sy in the Direction ad ; then the Particles of this Ray, when they arrive at the Space of Attraction at d , will be attracted by the new Force of the Medium sy in a Direction dg perpendicular to the Surface st ; and the Particles there being urged in two different Directions, ad and dg , can't possibly proceed in either of them but in one that is compounded of both, which will be inclined towards the Perpendicular dg .

178. But because the attracting Force increases according to some Power of the decreasing Distance from the Surface st on each Side (166.) therefore the new Direction of the Ray cannot

be

be a strait Line but a Curve, as it is continually urged in a perpendicular Direction by a Force uniformly accelerating the same. Wherefore during the Time of its being within the Space of Attraction $q w$, it must describe some Curve from d to k .

179. When the Particles arrive at k in the Limit of the Space of Attraction $v w$, they will then be again affected with an equal Force on every Side, and their Motion will now be again uniform, and they will proceed in the right Line $k p$, which is a Tangent to the Curve $d k$ in its lowest Point k : Consequently the Ray $k p$ must make Angle with, or be inclined to the incident Ray $a d$, which is also a Tangent to the Curve in its first Point d . All this follows from the Principles of *Mechanics*.

180. And since this is the Case of any other Rays $b e$, $c f$, &c. it is evident the whole incident Beam of Light $a c f d$ will pass, or be refracted in a curvilinear Direction thro' the Space of Attraction into the Beam $k p m n$, whose Velocity will now be greater than in the first Medium, by all that it acquired from the *accelerating Force* of this new *Medium*.

181. This Space of Attraction is however so very small, as not to be sensible in physical Matters; and may therefore be represented by a Right Line interceding the two Mediums.

182. Let therefore AB be the Surface of any dense *Medium*, as Glass, Water, &c. and suppose IC a Ray of Light incident upon it from the *Air* in the Point C . On that Point, as a Center, let a Circle be described, and let HE pass thro'

I

the

Fig. VII.

The DOCTRINE of REFRACTIONS.

the same Point C perpendicular to the Surface AB. Then is the Angle ICH called the *Angle of Incidence*, of which ab is the Sine.

183. The attracting Power of the Medium ABDG being greater than that of Air, will cause the Ray IC to be bent out of its first Direction ID into another CF, making a less Angle ECF with the Perpendicular EC; the Sine of which Angle is em , and is call'd the *Sine of Refraction*.

184. The Angle FCD measures the Quantity of the Refraction; and which we shall call the *Total Refraction*.

185. It is found by Experience, that the Sine of Incidence ab is to the Sine of Refraction em always in one constant Ratio, let the Angle of Incidence be what it will; which suppose to be as m to n : That is, it will always be as $ab : em :: m : n$.

186. If FC be a Ray passing out of a dense Medium into Air of a less refractive Power, it will then be refracted from the Perpendicular into the Direction CI; and the Sine of Incidence em will be to the Sine of Refraction ab as n to m .

187. If the Rays were all of *one Kind*, or equally affected by the attracting Force of Bodies, there would be an equal Refraction of all the Parts of a Beam of Light; but this is not the Case, for Experiments shew that *some Rays of the Beam will always be more refracted than others*. Thus, for instance, if IC be a Beam incident from the Air on any more powerful Medium ABDG, then Part of it will be refracted

Fig. VIII.

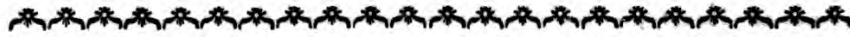
to

to r , making the Angle rCb of the *least Quantity*, and Cr is of Course the *least refrangible Ray* of that Beam.

188. Again, another Part of the Beam will be refracted to v , making the Angle vCb of the *greatest Quantity*; and therefore Cv is the *most refrangible Ray*. The Ray Cm , which equally divides the Angles vCr , is the *mean refrangible Ray*. And mCb is the Quantity of mean *Refraction*, and what for the future will be called *General Refraction* of the Medium; as the Angle vCr will be called the *Angle of Dissipation*, as being that into which the entire Beam IC is dissipated by Refraction.

189. The Rays of the dissipated Beam vCr are generally reckon'd of *seven Sorts*, because each Sort has a different Effect on the Fibres of the Optic Nerve, and excites a peculiar Idea of Colour in the Mind; and the dissipated Beam itself, if view'd as it comes from a Prism, appears variously colour'd from r to v . The Colours are in the followidg Order, *Red, Orange, Yellow, Green, Blue, Indigo, and Violet*.

190. Also dr , em , fv , are the Sines of the *least refrangible* or *red-making* Rays Cr , of the mean or green Rays Cm , and of the least refrangible or violet colour'd Rays Cv ; and when the Angle of Incidence ICH is small, the Angle of Refraction ECm is smaller; and in Case of such Angles as are less than 20 or 15 Degrees, *the Proportion of the Sines and of the Angles themselves may be reckon'd the same, without any sensible Error.*



C H A P. II.

The THEORY of the ABERRATION or DISSIPATION of heterogeneous LIGHT, explain'd and illustrated.

191. **T**HE next Step in our Business is to raise a few easy Theorems for ascertaining the Quantity of the *total and general Refraction*, and of the *Angle of Dissipation*, under all the Circumstances of *different Angles of Incidence*, and *Mediums of various Densities and specific Powers* of the Particles.

In order to this let the Quantities be represented by the following *Symbols, viz.*

m = The Sine of Incidence.

n = The Sine of Refraction.

D = Density of the Medium.

S = Specific Power of its Particles.

R = Refractive Power of the Medium.

T = Total Refraction.

A = Angle of Dissipation.

I = Angle of Incidence.

F = Angle of Refraction.

192. Then in small Angles, (190.) when the Refraction is made out of a less into a more powerful Medium, we shall have $m : n :: I : F$; and

and therefore $m : m-n :: I : I-F$; consequently $\frac{m-n}{m} \times I = T$, the Total Refraction, or Angle FCD, (Fig. VII.)

193. But if the Refraction be made out of a greater into a less refractive Medium, then $n : m :: I : F$, and $n : m-n :: I : F-I = T$; therefore $\frac{m-n}{n} \times I = T$, the total Refraction or Angle vCa.

194. Hence it appears, since n is ever as I , the Refraction T will always be as $m-n$, let the Ray be refracted which way it will. Also when I is given, $\frac{m-n}{n} = T$; or $\frac{m-n}{m} = T$, (192.)

195. Moreover since the refractive Power R of any Medium is as D when S is given, and as S when D is given, (170, 171) it will in general be as DS ; that is, we shall always have $R : DS$.

196. It is also evident, since Effects are always proportion'd to their Causes, that therefore, in a given Angle of Incidence, the Refraction T will be as the Cause R which produces it, viz. $T : R$; whence $\frac{m-n}{n} : DS : R : T$.

197. Lastly, it is plain that the Quantity or Angle of Diffipation vCr is some Part of the Refraction mCb ; or that $A = \frac{I}{p} \times T$, or $pA = T = \frac{m-n}{n} \times I$.

198. Let us now compare these Quantities in any two different Mediums ABD , which we will denote by X , and ABF , which let be call'd Y ; and then suppose the Quantities denoted by the
Roman

Fig. VIII.

The THEORY of DISSIPATION,

Roman Capitals above belong to X; let m , n , d , s , r , t , a , p , i , in small Italics signify the same Things in the Medium Y. And because we shall more generally consider the Refraction as made out of a greater X into a less refractive Medium Y, we shall use the Expression $\frac{m-n}{n}$ in our Comparisons. (See 194.)

199. And therefore we have for the first Analogy $\frac{m-n}{n} \times I : \frac{m-n}{n} \times i :: T : t$.

200. And when I and i are given, we have $\frac{m-n}{n} : \frac{m-n}{n} :: T : t :: R : r :: DS : ds$, (195.)

201. But to shorten the Work, let $\frac{m-n}{n} = b$, and $\frac{m-n}{n} = b$. Then $pA :: pa :: T : t :: bI : bi$, (199.)

202. These Analogies will afford us Theorems sufficient for our Purpose, after we have first determined the true Value of these several Quantities in such Mediums as will be necessary to consider, in regard to Vision, by refracted Light in dioptric Instruments, and particularly TELESCOPES and MICROSCOPES.





C H A P. III.

Of the General and Specific REFRACTION, and DENSITY of Mediums ; with the ANGLES of DISSIPATION, and THEOREMS for computing the same.

203. **T**HE Quantities to be determined, and their Values in different Mediums ascertain'd, are the following: (1.) The Ratio of m to n , or of the Sine of Incidence to that of Refraction. (2.) The refractive Power of the Medium. (3.) The Density of the Medium. (4.) The specific Power of its Particles. (5.) And the Angle of Dissipation.

204. The *First* of these is to be done very accurately, either by a *Prism* or a *Lens*, or many other Ways well known. The *Second* is determined by measuring nicely the Quantity of Refraction made in equal Angles of Incidence in several Mediums. The *Third* is known from the specific Gravity of the Bodies or Mediums taken by the *Hydrostatic Ballance*. The *Fourth* is found by dividing the refractive Power by the Density: And the *Fifth* is easily known by measuring accurately the Length of the colour'd Image of the Sun form'd in the Focus of a Lens placed by the Prism, for then having the focal Distance of the Lens, the Angle of Dissipation is known, and from thence also what Part it makes of the whole Refraction.

205. Most of these Quantities have been determined by Sir *Isaac Newton* for many different Mediums and Bodies, which I have transcribed from him as in the following Table.

The refracting Bodies.	The Proportion of the Sines of Incidence and Refraction of yellow Light.	The refracting Force of the Body.	The Density and specific Gravity of the Body.	The specific refractive Power of the Body.
	$m : n$	D.S. : R	D	S
A Pseudo-Topaz.	23 to 14	1,699	4,127	0,3979
Air.	3201 to 3200	0,000625	0,0012	0,5208
Glass of Antimony.	17 to 9	2,568	5,28	0,4864
A Selenite.	161 to 41	1,213	2,252	0,5386
Glass vulgar.	31 to 20	1,4025	2,58	0,5436
Crystal of the Rock.	25 to 16	1,445	2,65	0,5450
Island Crystal.	5 to 3	1,778	2,72	0,6536
Sal Gemmae.	17 to 11	1,388	2,143	0,6477
Alum.	35 to 24	1,1267	1,714	0,6570
Borax.	22 to 15	1,1511	1,714	0,6716
Niter.	32 to 21	1,345	1,9	0,7079
Dantzic Vitriol.	303 to 200	1,295	1,715	0,7551
Oil of Vitriol.	10 to 7	1,041	1,7	0,6124
Rain Water.	529 to 396	0,7845	1,	0,7845
Gum Arabick.	31 to 21	1,179	1,375	0,8574
Spirit of Wine well rectified.	100 to 73	0,8765	0,866	1,0121
Camphire.	3 to 2	1,25	0,996	1,2551
Oil Olive.	22 to 15	1,1511	0,913	1,2607
Lintseed Oli.	40 to 27	1,1948	0,932	1,2819
Spirit of Turpentine	25 to 17	1,1626	0,874	1,3222
Amber.	14 to 9	1,42	1,04	1,3654
A Diamond.	100 to 41	4,949	3,4	1,4556

206. In this Table the Ratio of m to n are as in the first Column; thus, for the Medium of *Water*, we have $m : n :: 529 : 396$; for *Glass*, (such as he used) $m : n :: 31 : 20$; and so of the rest.

DENSITIES and Specific REFRACTIONS.

207. The *second Column* shews the Ratio of the refractive Powers R and r for different Media; thus, for *Glass and Water*, we have $R : r :: 1,4025 : 0,7845 :: DS : ds$; and, with respect to *Glass and Diamond*, it is $1,4025 : 4,949 : DS : ds$.

208. The *third Column* contains the Ratio of the Densities D, d , of those Bodies. Thus, for *Glass and Water*, $D : d :: 2,58 : 1$. Of *Glass and Diamond*, $D : d :: 2,58 : 3,4$.

209. In the *fourth Column* are contain'd Numbers expressing the specific Attraction S, s , of the Particles of those Bodies; as, for *Glass and Water*, we have $0,5436 : 0,7845 :: S : s :: \frac{R}{D} : \frac{r}{d}$.

From this Table we may observe,

210. First, That the *densest Medium has not always the greatest refractive Power*. Thus the Density of *Water* 1 is greater than that of *Spirit of Turpentine* 0,874; but its refractive Power 0,7845 is less than that of the *Spirit* 1,1026: And hence it will follow, that a Ray of Light IC passing out of a dense Medium X into a rarer Y , may be refracted from C to b towards the Perpendicular CH . This is easily confirm'd by Experiments.

211. Secondly, *That Light may pass out of one Medium into another of different Density, and yet suffer no Refraction*. For those Media may have the same refractive Powers, as *Borax* and *Olive Oil* have each of them 1,1511, but different Densities 1,714 and 0,913. Consequently a Beam of Light IC passing out of a Prism of *Borax* X

K

into

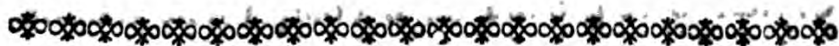
Plate III.
Fig. IX.

into another of *Oil Olive Y*, will not be refracted but pass straight on to *b*.

212. Thirdly, *Light may be refracted in passing out of one Medium into another, tho' the Density of each be the same*; for one may exceed the other in the specific Refraction of its Particles. Thus *X* may be a Prism of *Dantzic Vitriol*, and *Y* a Prism of *Alum*, whose Densities are the same, but their specific Refractions as 0,7551 and 0,657, and therefore a Beam of Light *IC* passing out of *X* into *Y* will be refracted from the Perpendicular *CH* into the Beam *Ca*.

213. Fourthly, *A Beam of Light may pass thro' two different Mediums, and be refracted from the last without any Dissipation of its Rays; and yet the Direction of the refracted Ray shall not be parallel to, but shall make an Angle with the Direction of the incident one.* This follows from the same Principles. But as it is the fundamental Article on which the Business of preventing the Imperfection of Vision, caused by the Error arising from the *Dissipation or different Refrangibility of Light* in dioptric Instruments entirely depends, it will be necessary to give a more ample and particular Elucidation thereof in the next Chapter.





C H A P. IV.

The DENSITIES, and general and specific REFRACTIVE POWERS of different Sorts of GLASS, CRYSTAL, &c. determin'd by EXPERIMENTS.

214. **T**HE Refraction of Light thro' any other Mediums than those which are used in dioptric Instruments, will be of no Use to consider; and since only the different Sorts of *Glass* and *Crystals* are fit for such Purposes, we must next enquire, what Effect they will have in refracting the Rays of Light, both *singly* and *combined* under all the different Circumstances of *Density*, *refractive Power*, *specific Refraction*, and *Dispation of the Rays*; first in the Form of PRISMS; and secondly, in the Form of Analogous LENSES.

215. For this Purpose it became necessary in the first Place to determine the Ratio of m to n , the several Mediums of *Glass* and *Crystal* as are in common Use, which I did by the following Method. In respect of any given Lens, double and equally convex, if we put r = Radius; d = Distance of the Radiant, or Object; f = to its proper focal Distance, and $b = \frac{m-n}{n}$; then by a Theorem demonstrated in my *Philos. Britannica*, we have this other Theorem, viz. $\frac{dr+rf}{2df} = b$.

The THEORY of Specific REFRACTION,

216. To get the Values of these Quantities, I caused Lenses of the several Sorts of Glass and Crystal to be ground on a Tool, whose Radius was just 21,5 Inches, so that in all of them $r=21,5$.

217: Then at the Distance of 436 Inches from a luminous Object, I placed a Screen in a dark Place, to receive the Image thereof, by the Lens moved forwards and backwards on a Rule divided into Inches and Tenths, till the Image became perfect; and then by Inspection the Value of f was known for the several Lenses, which subtracted from the whole Distance 436, left the Value of d , or the Distance of the Radiant from the Lenses severally.

218. Thus in a Lens of *white Flint-Glass* $r=21,5$, $f=18,75$, and $d=417,25$; therefore by the Theorem (215.) $\frac{dr+rf}{2df} = b = 0,5998 = \frac{m-n}{n}$; whence by putting $n=1$, we have $m-1 = 0,5998$; and so $m=1,5998$; therefore $m:n :: 1,5998:1 :: 1,6:1$ nearly.

219. If the Radiant be the Sun, then d is infinite, and the Theorem becomes $\frac{r}{2b} = 17,923$ for the *solar focal Distance*, which therefore is less than the Radius by 3,577 Inches.

220. In a Lens of *common Plate or Coach Glass*, I found $f=19,65$, and $d=416,35$, and $r=21,5$ as before; then $\frac{dr+rf}{2df} = 0,573$: whence (d being infinite) $\frac{21,5}{2 \times 0,573} = 18,76$ the Distance

of

and Specific GRAVITIES by EXPERIMENTS.
of the solar Focus. And in this Case $m : n ::$
 $1,573 : 1$.

221. In a Lens of *Crown-Glass* I found $f =$
 $21,25$, $d = 414,75$, and $r = 21,5$; therefore
 $\frac{dr + rf}{2df} = 0,5318$; the Distance of the solar Focus
 $= 20,214$; and the Ratio $m : n :: 1,5318 : 1$.

222. The *yellow Plate*, or *Venetian Glass*, as
also *Brazil Pebble*, have so nearly the same
Numbers for the focal Distance f , and the Ratio
of m to n with the *Crown-Glass*, that the Difference
is scarce discernable by Experiment, and there-
fore not worth Notice.

223. The next Thing to be done was to de-
termine the *Densities* of those several Mediums
or Lenses, which I did by taking their *specific*
Gravities by the hydrostatic Balance, and found
them as below.

The specific Gravity of	{	White Flint	— —	1 to 3,29
		Common Plate Glass	— —	1 — 2,76
		Crown Glass	— —	1 — 2,52
		Yellow Plate	— —	1 — 2,52
		Brazil Pebble	— —	1 — 2,62

224. Hence it appears, that the refracting
Powers of the several Lenses were very nearly
as their specific Gravities, which are nearly equal
in *Crown Glass*, *yellow Glass*, and *Brazil Pebbles*;
as also in that Glass call'd *Dutch White*, and in
Cornish Diamond.

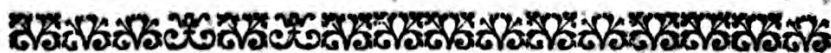
225. As *white Flint* and *Crown Glass* are in the
Extremes, and are most easy to be had for Op-
tical Uses, I shall chuse them for a Comparison.

And

Of the Prismatic REFRACTION,

And in these we have found $R : r :: DS : ds :: 0,5998 : 0,5318$ (see 196, 218, 221.) and also $D : d :: 3,29 : 2,52$; whence we have $S : s :: 0,5998 \times 2,52 : 0,5318 \times 3,29 :: 1,5115 : 1,7496$. Whence it appears, that tho' the *Density of Crown Glass* be less than that of *white Flint*, yet the *specific Attraction* of its *Particles* is considerably greater.

226. Having thus determined the *refractive Powers, Densities, and specific Attraction* of these two Sorts of Glass in general, we shall proceed to consider the *Phænomena* of Light refracted thro' them in the Form of Prisms, in regard to the Quantity of *Refraction, Dissipation, and Direction* of the incident and emergent Rays.



C H A P. V.

Of the Prismatic REFRACTION of LIGHT. THEOREMS determining the Angles of Total REFRACTION and DISSIPATION of RAYS; with the METHODS of rendering the refracted Beam Colourless, tho' inclined to the incident One in any given Angle.

Plate III.
Fig. IX.

227. **I**N treating of the Refraction of Light thro' a Prism, as BAD, it must be observed, that the Angle ICE contain'd between the Perpendiculars IC and EC to the Sides of the Prism, is always equal to the vertical Angle A contain'd under those Sides respectively; and this will always be the Case, whatever be the Quantity of

I of

of the Angle at D. But IC is here the incident Ray on the Side AB, and therefore the Angle of Incidence is equal to the Angle of the Prism at A, which is call'd the *refracting Angle of the Prism*, because the Refraction occasion'd by the Prism is proportion'd to that Angle, in as much as it is in Proportion to the Angle of Incidence. For since the Ratio of m to n is constant, the greater m is, the greater n will be; and also the greater will be their Difference $m-n$, which is the Refraction we speak of.

228. Wherefore let ABC be a Prism of *Crown Glass*, a Beam of Light sp parallel to the Base AC, will be refracted thro' it to q , and from them a second Time refracted and dissipated into the Beam rqv , the mean Ray of which gq contains an Angle gqb , which is the total Refraction of the Prism, and which is express'd by the Theorem $t=pa=bi$, (201.)

Plate IV.
Fig. I.

229. Suppose BCD an equal Prism of *white Flint Glass* in an inverted Position with respect to the other, and SP a Beam of Light incident on it in the same Direction as before, it will be refracted from P to Q, and from Q it will be refracted and dissipated into the Beam RQV, the mean Ray of which will make an Angle $GQH = T = Ap = bI$, (201.)

Fig. II.

230. Let QF be drawn parallel to gq , and the Angle GQF will be equal to the Difference of Refraction in the two Prisms; and therefore when they are both placed together, (as in *Fig. III.*) the same incident Beam sp refracted thro' the first from p to q , and thro' the Second from q to Q will be from thence refracted and dissipated into the Beam VQR , the mean Ray of which

The QUANTITY of DISSIPATION,

which GQ will contain an Angle GDH with the incident Ray sH , equal to the above-mentioned Angle GQH , or Difference of Refraction in the two Prisms.

231. Whence in this Combination of the two Prisms, since the refracting Angles at C and B are equal, we have $I=i$; and therefore $pA : pa :: b : b$ (201.) Therefore $A : \frac{b}{p}$, and $a : \frac{b}{p}$; and consequently $A - a : \frac{b}{p} - \frac{b}{p}$, which therefore will be as the Angle GQH of Dissipation.

232. Also we have $p : p :: \frac{b}{A} : \frac{b}{a}$, and therefore if the Angles of Dissipation rqv and RQV be exactly measured in the equal Prisms, the Quantities p and p will be known. Let the Values of A and a in this one Case be denoted by Z and z ; then in the general Equation $pAbi = apbI$, when $I=i$ (as in the present Case) we substitute Z and z for A and a , and it becomes $p : p :: zb : Zb$.

233. If we suppose the Angle ABC encreased, or the Angle BCD diminished to a proper Degree, there will be no compound Refraction, but the emergent Ray QH will be parallel to the incident one sp . In this Case $T=t$, and $pA = pa$, (201.) and so $A : a :: p : p$; and consequently there will be a Dissipation of the emergent Beam, and the Object will appear *colour'd* tho' in its proper Place. In this Case $bI = bi$, and so $I : i :: b : b :: 0,5318 : 0,5998$. Or the refracting Angles will be *inversly* as the refracting Powers of the Prisms.

Estimated, and annihilated.

73.

234. The Angle DCB may be so far diminished, or the Angle ABC so much increased, that the Angle of Dissipation RQV shall quite vanish, or the Rays of the refracted Beam shall be all parallel among themselves: Since in the general Theorem $pAbi = apbI$; if we put $A = a$, we have $I : i :: pb : pb :: zbb : Zbb :: z : Z$. Consequently when the refracting Angles of the Prisms are inversely as the Values of Z and z , (232.) there will be no Dissipation of Rays, and the Object will be seen by refracted Light without Colours.

Fig. III.

235. But in this very Case there will be a Refraction of the Beam from the Direction of its Incidence; for since $pA : pa :: T : t$, when $A = a$, or $A - a = 0$; then $T : t :: p : p$. And so $t - T$, or the compound Refraction by both the Prisms will be as $p - p$, a constant Quantity.

236. This may be effected two Ways, viz. by a Combination of two Prisms, or three; for if with a Prism of Crown Glass ABC you combine one of white Flint ECB, whose Angle ECB is to ABC as z is to Z , it will answer the Purpose.

Fig. III.

237. Also, if while the Prisms ABC and DCB remain together with equal refracting Angles, a third Prism of Crown Glass be added to them, as CDE, whose refracting Angle D, together with B shall be to the Angle of the Flint Prism C as Z to z the same Thing is effected; for then the incident Beam sp will be refracted thro' the three Prisms in such a Manner, that the refracted Beam PS will make an Angle SPH with the incident Beam sp continued, and with-

Fig. IV. }
V. }

Fig. V

L

out

out any Dissipation of the Rays, so that the Object or Radiant will now appear without Colours at G elevated above its true Place by the Angle sPG .

238. There are other Ways of shewing how this Effect is produced from such a Combination of two or three Prisms, but what I have said is sufficient to give the intelligent Reader an Idea of it, as it may most naturally be shewn by Prisms: I now proceed to consider the same Thing in a *Combination of Lenses in Telescopes.*



C H A P. VI.

Of the DOCTRINE of REFRACTIONS by LENSES of different Mediums, and the CONSTRUCTION of TELESCOPES with compound LENSES for preventing the DISSIPATION of Rays producing colour'd Light.

Plate IV.
Fig. VI.

239. **W**ITH regard to the Refraction of Light thro' Prisms and Lenses, it will appear to be the same in both when the following Particulars are consider'd. (1.) That a *Plano-Convex Lens* is only a Segment BLDC of a Sphere DBE, whose Radius is NL or NB.
240. (2.) That a Ray of Light SB, parallel to the Axis of the Lens NG incident upon it in the extreme Part B, will be refracted to some Part of the Axis. (3.) Let AB touch the Point B; then (4.) ABC will be a *Prism*, whose Base AC coincides with the Axis of the analogous Lens.
241. (5.)

How prevented by COMPOUND LENSES. 75

- (5.) Since the Ray of Light SB falls on a Point B, which is the same in the Lens and the Prism, it will be refracted in the same Manner from both; therefore (6.) the Ray will be not only refracted, but dissipated by the Lens to several Parts of the Axis between V, G, R. (7.) The refracting Angle of the Prism ABC is equal to the Angle of Incidence SBN, since they are severally equal to the Semi-angle of the Lens BNL. (8.) Continue NB at Pleasure, and in the Focus G of the mean refrangible Ray BG let a Perpendicular GI be raised to intersect NB produced in I, and SB continued in H. (9.) Then IBH is the Angle of Incidence, and (10.) then also HF, HG, HO are as the Quantities of the least, mean, and greatest Refraction of Rays in the Beam SB. (11.) The focal Distance LG will be less as the Angle of the Prism ABC, or Semi-aperture of the Lens LB, is enlarged. (12.) Let the Angle of Dissipation FBO be denoted by Z, and belong to a Lens of white Flint. (13.) Let f B o be the Angle of Dissipation belonging to a Lens of Crown Glass, and be denoted by z. (14.) The Sine of Incidence is to that of Refraction, as BG to NG; or $BG : NG :: n : m :: 1 : m :: 1 : 1,5998$. Also $Bg : Ng :: 1 : 1,5318$. (15.) FO is the Diameter of a Circle, into which all the Rays of a dissipated Beam or Pencil will be diffused, and therefore may be call'd the Circle of Dissipation or Aberration of Rays, arising from the different Refrangibility. (16.) When LG is large, we have $IH : IG :: n : m$, whence $HG = 0,5998 = b$. (17.) Whence $FG : HG :: \frac{1}{2}Z : 0,5998$,

or $FO : (2 HG =) BD :: Z : 1,1996 :: FBO : BGD$. And $z : 1,0636 :: f B o : B g D :: f o : BD$.

255. (18.) When B is extremely near to L, then $LG : NG :: 1 : 1,5998 \quad 1 : 1,6$ (nearly) $:: f : r + f$; whence $1,6 f = r + f$, and so $0,6 f = b f = r$, there-

fore $f = \frac{r}{0,6} = \frac{r}{\frac{3}{5}} = \frac{5}{3} r$ focal Distance of parallel Rays near the Axis in a Plano-Convex Lens. (19.) If the Lens be a double and equally Convex one as BPDF, then the Prism ABQ, being double of ABC, the Refraction will be HBG twice as much as before, consequently the focal Distance LG

but half LG; that is, in this Case $f = \frac{r}{2 b}$ as before was shewn from the general Theorem, (219.)

257. (20.) Since while BC remains the same, if the Angle ABC of the Prism becomes *greater or lesser* the Perpendicular NB will be *lesser or greater*, it is plain that the said Perpendicular of the Prism or Radius of the corresponding Lens will ever be *inversely as the refracting Angle* of the Prism. (21.)

258. As a Convex Lens BD answers in its Effects to two Prisms ABQ and ADQ placed together by their common Base AQ, so a Concave Lens answers to two equal Prisms put together on their Vertices or angular Points.

359. These Particulars connect the *similar Nature* of Prisms and Lenses, and plainly indicate what Lenses must be chosen for a Combination, that shall produce the same Effects as a given Combination of Prisms, and therefore what they must be that shall answer to that particular Combination of Prisms, (in Fig. V.) which transmits the Rays without Dissipation, and consequently such

such Lenses as will form the Images of Objects in the *compound Focus*, without the *Colours* as usual in Refraction by a *single Lens*.

260. If the Angle B of the Prism ABC of *Crown Glass* be equal to the Angle C of the Prism BCD of *white Flint*, then it is known by Experiment that the Angle D of the Prism CDE of *Crown Glass* must be very nearly equal to Half the Angle B of the first Prism, that the three Prisms together may refract the Rays of Light without Dissipation; and therefore the Sum of the two refracting Angles B + D is to the Angle C, as 3 to 2. Or $B + D : C :: Z : z :: 3 : 2$.

Fig. V.

261. Therefore if we take a double Convex Lens ABCG of *Crown Glass*, a double and equally Concave Lens CGAFD of *white Flint*, and a *Plano Convex Lens* DGFE of *Crown Glass*, and in a Position similar to that of the three Prisms, they will together make a *compound Object-Glass* of a Telescope, which will form the Images of Objects *without Colours*, and as distinctly as the Nature of such compound Refraction will admit of.

Fig. VII.

262. If the Telescope be short as those of the *Galilean Sort*, usually call'd *Opera-Glasses*, then the Convexity of the Lenses will be considerable, and in such Cases it is convenient to use three Lenses as above described, and as is represented in *Fig. VII*.

263. But in Telescopes of any Length exceeding 9 or 12 Inches, two Lenses are sufficient. For if the refracting Angle of the Prism ABC

Fig. III.

of

The DISSIPATION of RAYS prevented.

of *Crown Glass*, conjoin'd with a Prism ECB of *white Flint*, whose Angle $C = \frac{2}{3}B$, will refract the Rays without Dissipation (235.) it is plain, if we take any double and equally Convex Lens ABCG of *Crown Glass*, and another double Concave EBCGH of *white Flint*, these two Lenses put together will do the same Thing, that is, they will refract the Rays to their Axis without *Dissipation*, provided the Radii of the Concave have their *proper Proportion* to each other.

Fig. VIII.

Fig. III.

264. To determine which, let CG be a Perpendicular on the Base EB of the Prism BCE; then will it divide the said Base EB, or Angle ECB into two Parts ECG and GCB, which will be to each other as 1 to 3; consequently the Radius of the external Concavity in the Lens CD must be to that of the internal Concavity in the inverse Proportion, viz. of 3 to 1, (257.) But since the Angle $GCB = \frac{1}{2}ABC$, the Radius of the internal Cavity is the same as the Radius of Convexity in the Lens AB.

265. And here I must observe, that the Prism ECB must have the same Position (or be contiguous) to the Prism ABC, as represented in the Figure, to produce the same Effect; for if it be placed so that either of its Sides be perpendicular to the Base AC, tho' the two Prisms are in contact at C, there will be a *Dissipation of the Rays*, and the Object will appear *colour'd*.

Fig. VIII.

Fig. VII.

266. Hence appears the Reason why this Lens ECH should be a *double Concave* of unequal Radii: And consequently we see the true Form of the *compound Object-Glass* in a Telescope made

on this Principle, and constructed according to the *Theory*, and which is different from any Thing we have yet seen made.

267. What relates to the Number and Disposition of the Eye-Glasses, is the same as has been already observed (144.) in others; for here all the Care has been to form the Image *IM* without Colours, or *Aberration of Rays* from their *different Refrangibility*.



C H A P. VII.

THEOREMS for finding the focal Distances of compound LENSES, for any determined Lengths of TELESCOPES, OPERA'S, &c.

268. **I**T will be necessary to direct the young *Optical Analyst* how he is to proceed in finding the focal Distances of compound Lenses, or such as consist of two or three single ones of different Forms and refractive Powers; and to illustrate the Method by a Process for some of the most useful Cases.

269. And first let the compound Lens consist of three single ones, which we shall denote by the Letters *A*, *B*, and *C*; and the refractive Powers of the Mediums of which they are made be express'd by the small Letters *a*, *b*, *c*. Then for a double Convex Lens of unequal Radii *r*, *r*, and for any given Distance *d* of the Radiant,

we

we find the focal Distance f by this Theorem

$$\frac{p \, d r r}{d r + d r - p r r} = f. \quad *$$

270. Here $p = \frac{n}{m-n}$ and the refracting Power of any *Medium* being as $\frac{m-n}{n}$, or as $m-1$, (when $n=1$) as we have shewn (192); therefore a , b , and c , will be as $\frac{m-n}{n}$ in their respective *Media*, and so p will be as $\frac{1}{a}$, $\frac{1}{b}$, $\frac{1}{c}$; and the above general Theorem will be for a Lens of each Medium thus express'd :

$$\text{For A, } \frac{d r r}{a d r + a d r - r r} = f. \quad \text{For B, } \frac{d r r}{b d r + b d r - r r} = f.$$

$$\text{For C, } \frac{d r r}{c d r + c d r - r r} = f.$$

271. Now it has been shewn, that in the triple compound Lens, the first single Lens A is a double and equally convex one, and consequently the Theorem will (since $r=r$) stand thus for diverging Rays, $\frac{d r}{2 a d - r} = f$.

272. If one of the Radii (r) be infinite, then the said Lens becomes a *Plano-Convex*, and the Theorem for it is $\frac{d r}{a d - r} = f$. And such is the other outside Glass C of the triple Lens, as shewn (260.)

273. If the Radii (r_1, r_2) be both Negative, the Lens becomes a double and unequally Concave one, viz. $\frac{d r r}{-a d r - a d r - r r} = f$. 274.

* See my *System of Optics*, Page 60 and 76. Also *Philosophia Britannica*, second Edition; and the Table of *Theorems* which hereafter follow.

To double and triple Compound LENSES.

274. If one Radius (r) be infinite, then the Lens is a *Plano-Concave*, and the Theorem is

$$\frac{dr}{-ad-r} = f.$$

275. When both the Radii are equal, (*viz.* $r=r$) then the Lens is double and equally Concave, and its Theorem is

$$\frac{dr}{-2adr-r} = f.$$

And this we have shewn (261.) is the Form of the Middle Glass B in the triple compound Lens.

276. Since the Object Lens D is a *Plano-*

Convex, its focal Distance will be $\frac{dr}{cd-r} = f$,

and this will express also the Radiant Distance (d) with respect to the middle Lens B, whose

Theorem in the present Case being $\frac{dr}{-2bd-r} = f$,

for diverging Rays; it will be for converging

Rays $\frac{-dr}{2bd-r} = f$, for so the Rays will fall on

the second Lens, and therefore (d) will be negative.

277. Hence we have $d = \frac{fr}{2bf+r}$; and there-

fore $\frac{dr}{cd-r} = \frac{fr}{2bf+r}$ which will give $\frac{dr}{cd-2bd-r}$

$= f$, the compound focal Distance of both the Glasses C and B together.

278. Now since the middle Glass B is supposed to have the greatest refractive Power, or (b) greater than (c), therefore f last found will be negative; and as that is the radiant Distance, with respect to the third Lens A, or that next the Eye, the Rays will proceed from the second Lens B *diverging* upon it; and therefore the

M

Theorem

Theorem (in *Art.* 270.) will give $\frac{rf}{2af-r} =$ Ra-
diant Distance $= \frac{dr}{g d-r}$, by putting $c-2b=g$.

(277.) And from hence we get $\frac{dr}{2ad-gd+r} = f$, the
compound focal Distance of all the three Lenses
A, B, C, as required.

279. I have supposed the Lens C a Plano-
Convex, as it is the only genuine Form, (by
260.) and as the Theorems are more simple on
that Account; but if we had supposed it a
double and unequally Convex, we should in a
more tedious Manner have arrived at the same
Conclusion or *Formula* for the triple Lens;
for then we should have had this Theorem,

$$\frac{drR}{2adR+2bdR+rR-dcR-dcr} = f$$
; where by
supposing $2b-c=-g$, and R infinite, we get
$$\frac{dr}{2ad-gd+r} = f$$
, the same as before.

280. If this compound Lens be exposed to
parallel Rays, then d is infinite, and the Theo-
rem becomes $\frac{r}{2a-g} = f$, which is the Case of a
Telescope made with such a triple compound
Object-Glass.

281. When we suppose the Lens C to be a
Plano-Convex one, it is to adapt the Theorem to
the present Construction, where all the three
Lenses are Glass, and two of them A and C of the
same Sort of Glass; and consequently in this Case

$a=c$, and the Theorem $\frac{r}{2a-g} = \frac{r}{2a-2b+a} =$
 $\frac{r}{3a-2b} = f$.

282. We have shewn (218, 221.) that in white Flint and Crown Glass we have $a=0,5328$, and $b=0,5998$; but we may take $a=0,53$, and $b=0,6$, and be near enough the Truth for Use.

Thus for Example; let $r=10$, then $\frac{r}{3a-2b} = \frac{10}{1,59-1,2} = \frac{10}{0,39} = 2,82 = f$, the focal Distance of such a compound Object-Glass.

283. Suppose the compound Lens were to consist of two Glasses only, A and B; then A will be a double and equally Convex Lens of *Crown Glass, yellow Plate, or Pebble*, (224.) and B a double and unequal Concave, as we have shewn (263.)

284. The above Theorem for such a Lens becomes $\frac{drr}{-bdr-bdr-rr} = f$; and supposing (d) infinite, then $\frac{rr}{-rb-rb} = f$, the focal Distance of parallel Rays.

285. Now this is the Radiant Distance with regard to the interior Lens A; and therefore we have

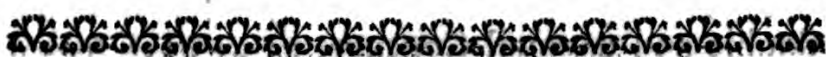
$$\frac{rr}{-rb-rb} = d = \frac{rf}{2af-r}; \text{ and putting } r+r=g,$$

we get $\frac{rr}{2ar-gb} = f$, the compound focal Distance of those two Lenses A and B.

286. But it has been shewn (264.) that in case of two such Lenses for taking away Colours, we shall have the Proportion of the Radii as 3 to 1; that is, $r:r::3:1$; consequently $r+r=4=g$, hence the Theorem becomes $\frac{3}{6a-4b} = f$

$= \frac{3}{3,18-2,4} = \frac{3}{0,78} = 3,84$ the focal Distance of both, when the Radius (r) of the double Convex A is 1.

287. By these Theorems the focal Distances for all compound Lenses of two or three different Mediums may be found, and thereby the Method of adapting them to Telescopes and Microscopes render'd universal, I might enlarge much on this Subject, if it were worth while, but what I have said is more than sufficient for the intelligent Optician.



C H A P. VIII.

*Remarks on the Nature, Construction, and Effects of the foregoing Achromatic * TELESCOPE.*

288. **T**HE Telescope in *Fig. VIII.* is such, as I think provides against all the Errors of Vision by refracted Light, in the best Manner that any one of this Kind can do; but it is at the same Time a *Memento*, that we ought to expect *no absolute Perfection in the Works of Art*; that is the sole Prerogative of the divine Being, and to be sought for in the *Works of Nature* only.

Fig. VIII.

289. There has been no Telescope made of the Construction here represented, that I know of:
The

* From the Greek *a privative*, and *χρῶμα*, *Colour*. But to say the Truth, no Telescope can shew an Object *without Colours*, if the Object be viewed with a *Lens*; for the ocular Lens only will cause a colour'd Image on the *Retina* in the Fund of the Eye.

The *Opera-Glass* in Fig. VII. has been made with a triple Object-Glass, but they are put together in a Manner different from that in which they here stand, and which we have shewn to be analogous to the three Prisms.

290. A Telescope, has also been made with a double Object-Glass, consisting of a *double Convex Lens*, and a *Plano-Concave*, with its plain Side *outwards*; and five Glasses at the Eye-End all differently placed from those in any of the Telescopes here represented in these Plates.

291. Now with regard to this or any Sort of refracting Telescopes, there are many Imperfections inseparable from them in the Nature of Things, and which therefore all our Art will be in vain applied to elude: For

292. *First*, There is no such Thing as a perfect Medium to be found, or such as will refract the Light with a Regularity necessary for perfect Vision. *Glass* of any Kind (being an *artificial Mixture*) is far from being such a Medium; nay *Crystal* itself, the most perfect and natural of all transparent Substances, will admit of no perfect Refraction on many Accounts.

293. *Secondly*, If the Medium were perfect, there is yet no such Thing as a perfect *Politure* of Surface to be effected by Art, which yet is necessary for a true Refraction of Light, and perfect Vision by it.

294. *Thirdly*, The Light incident on a Lens is reflected externally and internally from both the Surfaces, and many Times internally from one Surface to the other, which must necessarily occasion a Confusion in the Refraction
in

in some Degree ; for this internal Reflection of Light in a Lens is so considerable, as to occasion a *secondary Focus* in the Lens, in which Objects may be magnified much more than in the *common Focus*, tho' this has pass'd hitherto unobserved.*

295. *Fourthly*, The greater the Number of Glasses, the greater the Confusion in Refraction, and *Obscurity* in the Object, by Loss of Light at each Surface by Reflection and Refraction.

296. *Fifthly*, There will ever be an Error from the Figure of the Lenses, as we have shewn; and we can proceed only *one Step* in correcting it, so that there will ever remain an Imperfection of Vision on that Account.

297. *Sixthly*, The Error arising from the different Refrangibility of Light admits only of a *palliative Cure* ; for I believe there are but few People who will pretend to make a Telescope *in FACT as perfect as it is in THEORY*.

298. *Seventhly*, Besides, unluckily for this Telescope, that promises so much Perfection, it *cannot mend one Fault without making another*, for to annihilate the Colours it must be associated with a *Concave Lens* ; and such a Conjunction of dissimilar Lenses very much encreases the Error from the Figure of them, as we have amply shewn, (from 127 to 137.)

299. *Eighthly*, The Aberration of Rays from the Figure is still farther encreased, by our being obliged to have a *Concave Lens* of such a Form and Position as the Theory forbids in regard to that

* I have often magnified Objects this new Way, by a Lens 3 or 4 Inches Diameter, and 12 or 18 Inches common focal Distance, but the secondary Focus not more than 2, or $2\frac{1}{2}$.

that very Point. A *Plano-Concave* being the best Form, and its Concave Side placed *outwards*, the best Position, (84—88.) But neither of these Particulars can be admitted in this Telescope, (266.)

300. *Nintbly*, The Image IM should be reversed, or the second Image OP form'd, by two Glasses only, *viz.* EF and GH, as we have shewn, (144.) A Lens interposed between the Image IM and double Object-Glass, appears to me absurd; for if the said Image be free from the *Errors of the Figure* and *different Refrangibility both*, (as we are told) then it is certain it can only do Harm*. Yet such a Glass I find in all the Telescopes I have seen of this Construction.

Fig. VIII.

301. *Tentbly*, The Image OP should be view'd with two Eye-Glasses IK, LN, as we have before shewn (140.) which however it is not in those I have seen, tho' constructed with *seven Lenses*.

302. From these and many other Considerations disadvantageous to refracting Telescopes, which the very best Constructions are liable to, it plainly appears they are far enough from Perfection, or from affording that clear and distinct Vision of Objects which the World has been persuaded to believe, and expect; but this will be farther evinced, when we have first shewn the Nature and Construction of a *reflecting Telescope*.

C H A P.

* If it be said—it *amplifies the Field of View*, it is granted; but then at the same Time, it will *unequally refract the Rays*, and, in some Degree, it *regenerates the Colours* that were before destroy'd.



C H A P. IX.

The NATURE, CONSTRUCTION, and IMPROVEMENT of the CATADIOPTRIC or common REFLECTING TELESCOPE explain'd.

303. **T**HE Nature and Construction of a *Catadioptric* or REFLECTING TELESCOPE I have largely explain'd in my *System of OPTICS*, as also in my *Philosophia Britannica*; the principal Particulars and Properties on which it depends are as follow.

304. *First*, That the *Angle of Incidence* is equal to the *Angle of Reflection* in every Sort of Light.

305. *Secondly*, That the Image of an Object is form'd by Rays reflected from a *Speculum* in one and the same Point of its *Axis*, provided those Rays are incident *very near the Vertex* or *Axis* of the *Speculum*,

306. *Thirdly*, That therefore such an Image is entirely free from that Confusion and Colour of Rays, by which it is impair'd in the Focus of a common Lens.

307. *Fourthly*, That with regard to those Rays which fall at a Distance from the *Vertex* on the *Speculum*, there will be an *Aberration of such Rays*, or they will not be all reflected to the same Point of the *Axis* 'tis true, but then the Error here by Reflection is but as 1 to $2\frac{1}{2}$, or as 2 to 5, nearly compared with that by *Refraction*, when the *Radius of the Sphere* is the same.

308.

Of the CATADIOPTRIC TELESCOPE.

89

308. *Fifthly*, Therefore the Area of a *Circle of Aberration*, (93.) or confused Representation of each Point in the Object, will scarcely be a *sixth Part* so great by *Reflection* as by *Refraction*, the Ratio being as 1 to 5,76.

309. *Sixthly*, When the *focal Distance* of the *Speculum* is equal to that of the *Lens*, the Error PQ by reflected Light, will be that by *Refraction* but as 1 to 16.

310. *Seventhly*, That therefore the *Area* of the *Circle of Aberration*, or Confusion of the Image, is here by the *Speculum* but a 256th Part of what it is by a refracting *Lens*.

311. *Eighthly*, That upon the whole the Error is every way so small by reflected Light, that it is altogether insensible in Experiments, and in the Construction of Instruments, particularly the *reflecting Telescope* *.

312. These Considerations first gave Rise to that noble Invention, for it was soon observ'd, that since the Image was form'd so very perfect by reflected Light, it might not only be inverted as in a Refractor, but also be magnified at the same Time by a small reflecting *Speculum* properly placed, and yet so as to shew the Object very distinctly.

213. Hence a double Power of magnifying an Object is acquired in the *reflecting Telescope*, and consequently the *great Length* of Refractors necessary for a great Power of magnifying, is thereby avoided; and also the Conveniency of viewing Objects in general, and especially the celestial Bodies is thereby render'd very great.

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214.

* All these Particulars are but so many Corollaries from the Propositions in 83 and 92.

Plate IV.
Fig. IX.

314. For Illustration, suppose AB, CD are parallel Rays which fall on the large polish'd *Speculum* BD, and by it reflected to its Focus F, where an Image LM is first form'd of the distant Object. This Image is so perfect, that being plac'd a little farther than the Focus (f) of the small *Speculum* E from it, the Rays of Light which constitute this Image will fall upon, and be so reflected from the small *Speculum* as to form a *second Image* \mathcal{M} in its proper Focus *F* belong to the Distance EF of the first Image; and this secondary Image being erect, is view'd by the Eye thro' the Plano-Convex HK.

315. This is the most simple Structure of the common *reflecting Telescope*, and its Power of magnifying is thus computed: Suppose $BF=9$ Inches, the *focal Distance* of the large *Speculum* BD; and $Ff=1\frac{1}{2}$ Inch, of the small *Speculum* EG. Then if EG were placed just at its focal Distance *Ff* from the Image IM, it would reflect the Rays parallel to the Eye; and the Image would appear larger than the Object in proportion of 6 to 1, that is to say, the Object is magnified *six Times by the large Speculum* only.

316. But now let the Image IM be consider'd as a small Object to be magnify'd by the small *Speculum* EG, (in the Manner of a *MICROSCOPE*) then the small *Speculum* being removed farther from IM than its focal Distance, *viz.* to the Distance EF, an Image will be form'd thereof at \mathcal{M} as much larger than IM, as the Distance EF is greater than the Distance HF, which let be as 8 to 1. Then is the Object magnified in the Ratio of 6×8 , or 48 to 1; or it will

will appear 48 Times nearer or larger in Diameter thro' the Telescope than to the naked Eye. But this Power of magnifying by Refraction would require a Length of Tube not less than 8 or 10 Feet in a Refractor.

317. I need not mention that only so much of the second Image $\mathcal{F}M$ can be seen as falls within the Tube or Part next the Eye, viz. $a b$, and that is to be duly circumscribed by a Hole in a Piece of Brass, call'd a *Diaphragm*, placed in the Focus of the Eye-Glass HK . Also at the End of the *Eye-piece* at (e) there ought to be a small Hole of a determinate Bigness, to shew the Object in its Image $\mathcal{F}M$ distinctly; if this Hole be too big or too little, the Telescope will be vitiated in proportion.

318. But this is not the Construction of the reflecting Telescope in common Use; for in that you always find a Glass Lens $(c d)$ placed just at the End of the Eye-Piece where it screws on to the large Tube. This Glass reduces the secondary Image $\mathcal{F}M$ to a lesser one im , (in the Manner as has been shewn in the Refractor (97, &c.) and so this Image im , which we view in the common Refractor, is made partly by the Speculum EG , and partly by the Lens cd ; and so is neither one Thing nor the other; that is to say, it is not form'd by *Refraction* or *Reflection* alone, but by both together; and makes the Instrument what it is truly call'd, a CATADIOPTRIC, instead of a CATOPTRIC TELESCOPE, as it ought to be.

319. In the *Catoptric Telescope* above described the Images are both form'd *entirely by refracted*

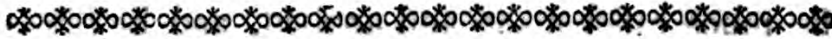
A true CATOPTRIC TELESCOPE described.

Light, and therefore not affected by the different Refrangibility of the Rays, consequently the Vision is as perfect as Art and Nature can make it by Instruments.

320. But in the common *Catadioptric Telescope*, the first Image only is form'd by Reflection, but the second Image is made by *Reflection and Refraction conjointly*; for the Rays reflected from the small Speculum EG fall converging on the Lens cd, and by it are refracted to f , short of the proper Focus F to which they tended before.

321. Now the Rays passing thro' the Lens cd must necessarily suffer a different Refraction, and consequently the Image im must be thereby render'd imperfect by the *Aberration of Light*, and the Colours and Confusion of Rays (by which it is form'd) consequent thereupon. Whence we may conclude, that this Glass being associated with the small Mirror does much Harm, and *lessens the Effect and Value of a reflecting Telescope.*

322. Besides it will appear that this Glass is not only injurious in itself in that Situation, but that it is injudiciously placed there, since the Image FM (form'd by *Reflection* only in the *Catoptric* or *true reflecting Telescope*) is to be view'd with the same Degree of Distinctness in the compound Focus of two Eye-Glasses placed together, as we have before directed, (140.) and as they are represented in the Figure by HK, and LN. And when the Glasses are thus disposed, I see nothing more at present that can add to the Perfection of Vision by such an Instrument *in the Form they are now made.*



C H A P. X.

The Imperfection of VISION by the Achromatic TELESCOPE shewn by many EXPERIMENTS, and the great SUPERIORITY of the REFLECTING TELESCOPE evinced.

323. **F**ROM what has been said of the Nature of *refracting and reflecting Telescopes*, it evidently appears, that if any refracting Telescope could truly correct the Errors arising from the *Figure of the Glasses*, and also from the *different Refrangibility of Rays*, it would be in its own Nature more perfect than the *reflecting Telescope*, because this latter is not free from the *Errors of the Figure of its Specula and Lenfes*.

324. And since we find the Reflector with this Imperfection so capable of a *double magnifying Power*, how much more might we expect it in that Refractor which pretends to be without any at all? If the *Reflector*, notwithstanding the Errors of the Figure, Loss of Light in the Metals, and other Disadvantages, will yet give a clear and distinct View of an Object by means of a *second Image*, who would not expect this noble Effect in a much higher Degree in a *Refractor* subject to no Sort of Errors, sustaining little Loss of Light, and having all the Perfections that the Science of Dioptrics can impart?

325. Any one would readily conclude such a Consequence must necessarily follow; and that now we should have *double Refractors* equally good,

good, if not superior to Reflectors, that should be much easier made and applied, and be purchased at a much less Expence. But how much must be abated of these great Expectations, will appear by attending to the following Experiments.

Fig. X.

326. I took the double Object-Glass of a Three-foot Refractor, $A B E H G$, whose focal Distance $A F$ was about 30 Inches, and combined it with the triple Object Lens $a b d e g$ of an *Opera*, whose focal Distance $e f$ was 6 Inches: These two compound Lenses were so adjusted as to make of the first small Image $L M$ a large one $\mathcal{Y} M$, after the same Manner by Refraction, as the same two Images were before form'd by Reflection.

327. This second Image I view'd in the best Manner by the two Eye-Glasses $H K$, LN , as in the Reflector of the same magnifying Power, but found a very dissimilar Effect. The Object in the double Refractor, instead of being bright and distinct as in the Reflector, was dark and confus'd; the Title Page of a Book at a Distance, clear and easily legible in one, was very obscure and unintelligible in the other.

328. It appears by this, which I look upon as an *experimentum Crucis*, that both the first and second Images in the *double Refractor* were far less perfect than in the *Reflector*; and tho' I gave it all the Advantages I could think of, it was to no Purpose, no Object could be view'd by it with any tolerable Degree of Distinctness.

329. As I knew the Glasses I had taken from the Telescope and Opera (306) were not precisely of the same Figure and Position the Theory required, I took the Trouble of making others that were so, but still found it an unsuccessful Attempt; and since, by Experience, I found so great a Difference in the Effects of a *double magnifying Power*, I was determined to try what I could learn from a Comparison of the single magnifying Powers of these two Instruments, which I perused in the following Manner.

330. I took the compound Object-Glass of the Opera, (306.) and form'd it into an *astronomical Telescope*, with a triple compound Eye-Glass, (see *Art. 39, 40, &c.*) The Instrument thus constructed magnified about six Times, and was truly an *achromatic Telescope*, or free from Colours. I then took a *Reflector* whose focal Length of the *larger Speculum* was *six Times* that of the small one, and therefore an Object viewed by these two Speculums only was magnified just as much as in the Refractor, and also without Colours; but there was no Comparison between the Distinctness of Vision in one and in the other. The Letters in the Title Page of a Book placed at a proper Distance, were extremely clear and well defined in one, that seem'd as it were in a misty Atmosphere, or whitish Medium in the other. I could easily read a Paragraph in the News Paper at a certain Distance with the *two Speculums*, but scarcely one Word of it with the *two compound Lenses* could be distinctly discern'd.

331. Having found so great a Difference in the Appearance of Images formed in the solar Focus of a compound Lens and Speculum, I thought it would be necessary to compare the Appearance of the Images form'd in the proper or conjugate Focusses of a small triple Lens and and Speculum, applied as an Object-Glass in a compound Microscope, (see *Art. 75. &c.*) which I did by giving the same Length to the Images, and viewing them with the same Eye-Glasses; and consequently the small Objects were equally magnified in the refracting and reflecting Microscope, but with how great a Difference in Point of Clearness and Perfection in all the Circumstances of Vision, those only will be able to conceive who shall try the Experiment themselves.

332. I resolved further to see what Difference there would appear in a Landscape form'd by a double *achromatic* Lens, and a Speculum of the same focal Distances; and having applied both in a *camera Obscura*, the Success was what I now expected. The Images of Objects in the Picture of the Lens were hazy, and towards the Margin indistinct; whereas those of the Speculum were excessively clear, luminous, vivid, and distinct, almost equally to the Borders of the Field, of equal Extent with the other.

333. There remain'd yet another Method of shewing the very great Difference in the Perfection of Images form'd by those compound heterogeneous Lenses, and the simple Reflection of Mirrors; and that was to apply them severally as SOLAR MEGALASCOPES: The Speculum and
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compound Lens were of the same focal Length, and the Object equally illumined and magnified in each ; but the Image of the former compound with that of the latter, was sufficient to shew how much the simple Operations of Nature excel those which are more compound, forced, and artificial.

334. I frequently observed in trying to find the Focus of a compound Lens, that it was attended with unusual Difficulty and Uncertainty ; I could scarcely tell when I had got it, or when not, so dubious was this Point, till after I had practised with them a considerable Time: And even now, unless the Sun be very clear, its Image is so indistinct in the *compound Focus*, that I cannot easily determine where it is most perfect ; whereas in a *single Lens* of the same focal Distance, the Image of the Sun at once appears distinct, in one Place only, and ever so little a Distance on one Side or the other, induces a sensible Confusion.

335. Another Thing discover'd itself in these Experiments about the Focusses of compound and single Lenses, which at first rais'd some Admiration ; it was this, the Subject on which I took the Focus of the Lenses, or Image of the Sun, was a Piece of Cork fix'd on to the End of a long Rule divided into Inches and Tenths ; the Surface of the Cork was made black with Ink ; and the Moment I had the Focus of the single Lens on the Cork, it began to smok intensely ; but when the Focus of the compound Lens was received on the same Cork, no Smoak appear'd in that sudden Manner as before, nor

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The Superiority of a REFLECTING TELESCOPE.

till after it had remain'd a sensible Time on the Cork, and then the Smoak and Burning were not comparable to that produced by a single Lens.

336. It appears by the two last (and some other) Experiments, that the solar Rays are more perfectly united in the Focus of a *single*, than in that of a *compound Lens*, and that their Force and Action is much greater. But since in the solar Focus of the same Diameter made by a Speculum, there can be no Error from the different Refrangibility of Rays; and that the Error of the Figure is here but a sixteenth Part so great as in the Lens, it must evidently follow, that the Perfection of the solar Image of the Speculum must greatly exceed that of the single Lens in all Respects, and much more that of the compound one.

337. Upon the Whole, therefore it appears, both by Theory and Experiments, *that a Combination of Lenses is far from rendering Vision perfect in refracting Telescopes; and that it is in every respect much inferior to the Effect of a Speculum in a REFLECTING TELESCOPE for exhibiting the best and most pleasing View of Objects.*





C H A P. XI.

The common THEOREMS of OPTICS tabulated, and rectified to express universally the focal Distance of RAYS reflected from MIRRORS, or refracted thro' LENSES of any Medium.

338. **A**S I know by too much Experience, how easy it is to make a Mistake amidst so great a Variety of Characters, affected with such Contrariety and Alternation of Signs, in the vast Multiplicity of *complex Theorems* necessary in all Optical Enquiries and Investigation; and that when any such Mistake is made, how widely it makes us stray from the Truth we are seeking, and often occasions our taking up with erroneous Positions, or totally baffling our most laborious and pertinacious Pursuits; I presume it will be very acceptable to the young Optician to have all the useful Theorems placed together in one View, as in the following synoptic Tables, where he will immediately find the Theorem that suits his Purpose in any optical Enquiry. Observing only that the Thickness of the Speculum or Lens is supposed to be inconsiderable in comparison of the Radius of Curvature; and then in CATOPTICS we have the following Theorems for the different Sorts of Rays.

339. SPECULUMS.

Rays.	Convex.	Concave.
Diverging	$\frac{dr}{2d+r} = f.$	$\frac{-dr}{2d-r} = f.$
Parallel	$\frac{r}{2} = f.$	$\frac{-r}{2} = f.$
Converging	$\frac{-dr}{-2d+r} = f.$	$\frac{dr}{-2d-r} = f.$

340. In DIOPTRICS the Theorems for determining the *Foci* of Rays by *simple Refractions* out of *Air into Glass*, whether *Convex* or *Concave*, and *vice versa*, are as in the Table annexed.

I. Out of AIR into GLASS.

Rays.	Convex.	Concave.
Diverging	$\frac{m d r}{m d - n d - n r} = f.$	$\frac{-m d r}{m d - n d + n r} = f.$
Parallel	$\frac{m r}{m - n} = f.$	$\frac{-m r}{m - n} = f.$
Converging	$\frac{m d r}{m d - n d + n r} = f.$	$\frac{-m d r}{m d - n d - n r} = f.$

341. II. Out of GLASS into AIR.

Rays.	Convex.	Concave.
Diverging.	$\frac{-n d r}{m d - n d + m r} = f.$	$\frac{n d r}{m d - n d - m r} = f.$
Parallel	$\frac{-n r}{m - n} = f.$	$\frac{n r}{m - n} = f.$
Converging	$\frac{-n d r}{m d - n d - m r} = f.$	$\frac{n d r}{m d - n d + m r} = f.$

342. As every Lens has two refracting Surfaces, there will be two *Radii* belonging to each, viz. r, r ; and because of a double Refraction, the Theorems for Lenses will be very complex, but may be abbreviated by putting $\frac{m-n}{n} = a$, or if $n=1$, then $a=m-n$. Whence the Theorems will be for the several Forms of Lenses as in the following Table.

343. I. A LENS with unequal *Radii*.

Rays.	Convex.	Concave.
Diverging	$\frac{d r r}{a d r + a d r - r r} = f.$	$\frac{d r r}{-a d r - a d r - r r} = f.$
Parallel	$\frac{r r}{a r + a r} = f.$	$\frac{r r}{-a r - a r} = f.$
Converging	$\frac{-d r r}{-a d r - a d r - r r} = f.$	$\frac{-d r r}{a d r + a d r - r r} = f.$

344. II. A LENS with equal Radii.

Rays.	Convex.	Concave.
Diverging	$\frac{dr}{2ad-r} = f.$	$\frac{-dr}{2ad+r} = f.$
Parallel	$\frac{r}{2a} = f.$	$\frac{-r}{2a} = f.$
Converging	$\frac{-dr}{-2ad-r} = f.$	$\frac{dr}{-2ad+r} = f.$

345. III. A LENS with one Radius (r) infinite.

Diverging	$\frac{dr}{a'd-r} = f.$	$\frac{-dr}{ad+r} = f.$
Parallel	$\frac{r}{a} = f.$	$\frac{-r}{a} = f.$
Converging	$\frac{-dr}{-ad-r} = f.$	$\frac{-dr}{ad-r} = f.$

346. IV. A LENS with one Radius (r) negative.

	Unequal.	Equal.
Diverging	$\frac{-dr r}{adr - adr + rr} = f.$	$-d = f.$
Parallel	$\frac{-r r}{ar - ar} = f.$	$\frac{-r r}{0} = f.$
Converging	$\frac{dr r}{adr - adr + rr} = f.$	$d = f.$

347. These are all the useful Theorems in Dioptrics for finding the *focal Distance* of any *Speculum* or *Lens*. And the Reader will observe they here stand in a different Form from what he finds them in other Books of Optics; the Reason is, because in vulgar *Optics*, and in practical Cases, it has been thought sufficient to take the Ratio of the Sines of Incidence and Refraction out of Air into Glass to be that of 3 to 2; and if that were true, we should have $\frac{m-n}{n} = \frac{1}{2}$ $= a$, or $2a = 1$; then, for Instance, in an equally Convex Lens the focal Distance would be $\frac{dr}{d-r}$; and for parallel Rays, $f=r$, as usually represented, and not as you find them in *Art.* (344.)

348. But this Position is very erroneous, and requires as much to be corrected, if not more, than those we have already consider'd: Because, in the *first* Place, 'tis an *Opprobrium* to the *Science* of OPTICS to have such a vulgar Error continued in it without Notice. *Secondly*, it supposes all the Sorts of Glass we use to have the same refractive Power, which is false and absurd: And *thirdly*, many late Inventions in Optics require a due Consideration of the real refractive Power and focal Distances of Lenses, and therefore make it necessary to have proper Theorems to represent and compute them by, and such I have here supplied in the above Tables.

349. And in order to fit them to every Sort of transparent Medium that can be of Use in Optics, I have nicely examined, and by Experiments determined the Ratios of the Sines of Incidence

105 Ratios of the SINES of Incidence and Refraction.

Incidence and Refraction in Water, Glafs, Pebble, &c. with respect to the *least, mean, and most refrangible Rays*, and have express'd them in the following Numbers.

350. In *New River Water* I find the Ratio of $m : n :: 1,3233 : 1$; which is so near the common Ratio $1,3333$ to 1 , that the Difference is not worth regarding. So that for the several Rays we have

For Red, $m : n :: 1,32727 : 1$ { and $a = 0,32727$
 Green, $m : n :: 1,33333 : 1$ { $a = 0,33333$
 Violet, $m : n :: 1,33939 : 1$ { $a = 0,33939$

351. In Pebble, Crystal, yellow Glafs, and Crown Glafs, I find the Ratios to be as follow.

For Red, $m : n :: 1,52213 : 1$ { and $a = 0,52213$
 Green, $m : n :: 1,53180 : 1$ { $a = 0,53180$
 Violet, $m : n :: 1,54147 : 1$ { $a = 0,54147$

352. The Glafs hitherto most in Use (though the worst of all on many Accounts) is *common Plate Glafs*; we have in this the following Analogies.

For Red, $m : n :: 1,563 : 1$ { and $a = 0,563$
 Green, $m : n :: 1,573 : 1$ { $a = 0,573$
 Violet, $m : n :: 1,583 : 1$ { $a = 0,583$

353. Lastly, in *white Flint Glafs*, from a Trial of many Sorts, and taking the Mean of all, the Numbers came out thus :

For Red, $m : n :: 1,5889 : 1$ { and $a = 0,5889$
 Green, $m : n :: 1,5998 : 1$ { $a = 0,5998$
 Violet, $m : n :: 1,6107 : 1$ { $a = 0,6107$

354. And thus the Values of m , n , and a , are known for LENSES, MIRRORS, and RAYS of every Sort; and by which any of the foregoing Theorems may be applied in optical Investigations, to express the Truth to the utmost Exactness that can be required. And it is very notorious, how faulty all our Computations must be that are made by Theorems not thus rectified; such Inaccuracies are inconsistent with the Precision of Truth, and the Honour of Science. The *Mechanical Optician* takes it for granted, that the Distance of the solar Focus is equal to the Radius, not knowing that in a Lens of 20 Inches only it is less by $3\frac{1}{2}$ Inches (219.) if it be of *white Flint*; and in the most common or *Plate Glass*, the Difference is near 3 Inches in 20, (220.)

355. Let the Radius of a Lens be of any Quantity, it appears from the foregoing Tables, (344.) that the focal Distance of parallel Rays is to the Radius as $r : \frac{r}{2a} :: 2a : 1$.

Or, R : F :: 1,1996 : 1 in a Lens of *white Flint*.

R : F :: 1,146 : 1 ——— of *Plate Glass*.

R : F :: 1,0636 : 1 ——— of *Crown ditto*.

356. Hence it appears, that if the Object-Glass of a Telescope be *Plate*, whose Radius of Convexity is $114\frac{1}{2}$ Inches, the Distance of the *solar Focus* will be but 100 Inches. Therefore in *measuring the Distance of an Object by a Telescope, or the Dimensions of the heavenly Bodies by the new Micrometer*, or

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The Use of those THEOREMS.

in the Solution of any *Problem depending on the Radius or geometrical Focus of the Object Lens*, very great Errors must ensue, if strict Regard be not had to those important Particulars, and how much this concerns not only the *Optician*, but the *Astronomer*, and every other *Artist* to whom the Science of Optics can be of Use, I may hereafter take another Opportunity to shew more at large.





C H A P. XII.

An Illustration of the METHOD invented by Mr. CALEB SMITH, for correcting the ERRORS arising from the different Refrangibility of RAYS, by Means of a Catadioptric SPECULUM and Concave LENS.

357. **T**HE first Discovery of the Art of correcting the Errors from the different Refrangibility of Rays, was exemplified in those Refractions which are made by a catadioptric SPECULUM and a LENS, properly associated together in their common Axis. And tho' the Principles on which it depends, are the same with those in the Method by Lenses already explain'd; yet as the Process is different, we shall, for the Sake of Variety, give the following brief Explication of it.

358. Let $A B C D$ be a Glass Speculum, foliated on the convex Side $C D$; put $r =$ Radius of the concave Surface $A B$, and $r =$ Radius of the convex one $C D$. Suppose $E A G C$ a Beam of Light parallel to the Axis $B N$, and incident on the *Speculum* in A , it will there be refracted to the second Surface at C , from whence it will be reflected to the first Surface again at O ; and there be a second Time refracted to the Axis, which

Plate III.
Fig. X.

The ERRORS of REFRACTION corrected,

the *most refrangible* Rays will intersect in F, and the *least Refrangible* in H, and FH will be the Aberration or Error occasion'd by the Diffipation of the Beam in the two Refractions.

359. Let n be the Sine of Incidence out of Glafs into Air; and m, m , the Sines of Refraction of the most and least refrangible Rays; then since the Beam first passes out of Air into *Concave Glafs*, we shall have the focal Distance $f = \frac{-mr}{m-n}$, for *Violet* Rays, (340.) put $m-n=a$, then $f = \frac{-mr}{a}$.

360. The Beam of Rays is now dissipated and refracted to the Surface CD at the Part C, where the Rays will now be reflected as from a concave Speculum; for which the Theorem (in vulgar Optics) is $\frac{-dr}{2-dr} = f$; but here $d=f$, in the foregoing Article, therefore we shall have $\frac{-mrr}{2mr-ar} = f$, the focal Distance of the reflected violet Rays.

361. These Rays are now to be consider'd as *converging*, and incident on the convex Surface of Air, OB, in the Point O. The focal Distance

of which will be $f = \frac{-ndr}{md-nd-mr} = \frac{-ndr}{ad-mr}$, (341.)

whence we get $d = \frac{mrf}{af+nr} = \frac{mrr}{2mr-ar}$; and

from thence, $f = \frac{nr}{2mr-2mr+2nr} = BF$, by restoring the Value of (a) .

362. In like Manner the focal Distance of the

the red Rays will be $\frac{n r r}{2 m r - 2 m r + 2 n r} = B H :$

Therefore the Value of the Aberration $B H -$

$$B F = F H = \frac{n r r}{2 m r - 2 m r + 2 n r} - \frac{n r r}{2 m r - m r + 2 n r} ;$$

this when reduced, and abbreviated, (by putting $r - r = g$, and $m - m = q$) will stand thus

$$\frac{g q n r r}{2 m m g^2 + 2 m n g r + 2 m n r g + 2 n^2 r^2} = F H.$$

363. If this Aberration be compared with the focal Distance of the Speculum, we shall have $F H : F B :: g q : g m + n r ;$ and,

putting $F H = A$, we have $A = \frac{f g q}{g m + n r}.$

364. By supposing $r = r$, we find $f = \frac{1}{2} r$, as in common Optics ; and then $A = 0$, or the Aberration is nothing in a Speculum, whose two Surfaces are parallel to each other.

365. If r be supposed infinite, or AB a plane Surface ; then the focal Distance is $f = \frac{n r}{2 m} ;$ and

the Aberration $A = \frac{m - m}{m} \times f ;$ which is in this Case a *Maximum*.

366. In this Respect the *Speculum* is only a *Plano-Convex Lens* foliated on its Convex Surface ; by which Contrivance or Artifice, a much shorter focal Distance and *less Dissipation* of Rays are produc'd by one refracting Surface, than could have been done in the same Glafs used as a Lens ; for in the Lens the focal Distance for parallel

Rays is $\frac{n r}{m - n} = f$, which is nearly twice the Ra-

dus ; whence in the Speculum $\frac{n r}{2 m} (= f)$ is but
about

The ERRORS of REFRACTION corrected,

about $\frac{1}{3}$ of the same Radius of Convexity; and therefore the focal Distance of the Speculum is but about $\frac{1}{3}$ Part so large as that of the Lens.

367. Then as to the *Aberration* of the most different heterogeneous Rays, *viz.* the *Red* and

Violet; for the *Red*, we have $f = \frac{nr}{m-n}$, and for

the *Violet*, $f = \frac{nr}{m-n}$, whence the *Aberration*

$\frac{nr}{m-n} - \frac{nr}{m-n} = \frac{m-m \times nr}{m-n \times m-n}$ in the Lens, is very

large if compared with that in the Speculum

$\frac{m-m}{m} \times f$, as will appear by reducing them both

to Numbers.

368. Mr. SMITH uses Sir ISAAC NEWTON'S Numbers, *viz.* $n=100$, $m=156$, and $m=154$;

therefore $m-m=n=2$, $m-n=56$, and $m-n=54$; whence the *Aberration* in the Speculum $A =$

$\frac{2f}{154} = \frac{2r}{462}$, but that of the Lens is $A = \frac{200r}{3024} =$

$\frac{1}{15}r$ nearly; whence the two Errors are as $\frac{1}{15}$

to $\frac{1}{231}$, or as 231 to 15.

369. Now it is known from the Doctrine of Refractions before explain'd, any Degree of Refraction or *Aberration* produced by a given Incidence

and Radius of Sphericity in a convex or concave Lens, will require a concave or convex Lens of the

same Radius and Angle of Incidence to correct that *Aberration*, and in that Case the Rays incident

on the first, and those which emerge from the second Lens, will have the same Direction; and

consequently an Association of two such Lenses can be of no Use in Optics, as we have shewn.

370. But if in any Lens we can by any Expedient produce, from a given Angle of Incidence, a Degree of Refraction or Aberration different from what is peculiar to the same Lens by itself, or any other Lens of the same Radius and Angle of Incidence; then two Lenses of a contrary Figure may be found with different Radii, and Angles of Incidence that shall have the same Aberration; and consequently that the Error of one may be corrected by the other; and at the same Time the incident and emergent Rays shall have a different Direction, or make a given Angle with each other. (See Chap. V.)

371. This we have shewn how to perform with two Lenses of a *different specific Refraction*; and here we have shewn, that by Means of Reflection, a Lens shall have a different Refraction and Aberration from that which is proper to it as a Lens only, with a given Incidence of Rays. Therefore there is some *concave Lens* I K L M, whose Aberration will be equal to that of the Speculum A B C D, and will correct it, or converge the Rays A H, O F to one Point N in the Axis.

372. Suppose the Radius of the Surface I K be denoted by x , and that of the Surface L M be called y : Then the general Theorem for a double concave Lens $\frac{ndxy}{-mdx + ndx - mdy + ndy - nxy}$
 $= f(342, 343.)$ And if we put $y = f$, the Theorem will be abridged to this Form $\frac{mdx}{nd - md - nx} = f = y = K N$;
 but as the Lens is here placed, it receives the Rays in a converging State, and therefore the Distance $d = F K$ is negative, and the Theorem becomes

becomes $\frac{mdx}{nd+nx-md} = f = KH$, for the least refrangible or red-making Rays OH ; and for the Violet OF , the Theorem is $\frac{mdx}{nd+nx-md} = KF$, whence $KH - KF = \frac{mdx}{nd+nx-md} - \frac{mdx}{nd+nx-md} = FH$, will be the Aberration of this concave Lens with regard to those incident Rays.

373. Now since the Aberration of this Lens must in this Case be equal to that of the Speculum, that so it may correct it, we must make an Equation of each, which, when properly reduced and order'd, will give $x = \frac{dd + dA \times m - m}{nA - m - m \times d}$ the Radius of the Surface IK next the Mirrour.

374. In this Equation the Quantities A , m , and n , are given; and d may be assumed at Pleasure of any Value between $\frac{m-n}{m-m} \times A$ and $\frac{mA}{m-m}$ (the Reason of which Limitation appears from the Process at large), and therefore x being known, the other Radius y is also known from the Equation in Article (372.)

375. The Point N thus determined is the compound Focus of the Speculum and Lens conjointly where the Images of distinct Objects are all form'd without Colours or Confusion, and may be view'd very distinctly by an Eye-Glass, if it could be conveniently applied for that Purpose.

376. But there lies the Difficulty; the Focus N must either be thrown by a small reflecting Plane

Plane to a Hole in the Side of the Tube or Telescope, as in the *Newtonian Form*; or else reflected back thro' a Hole in the central Part of the large Glase Speculum, by a small *Concave*, as in the *Gregorian* or common Form of a Reflector. But as the Trouble and Difficulty will be every way not a little, we believe it will not be attempted in Practice, especially as the chief End proposed is only to separate the two Images form'd by the first and second Surfaces of the Speculum farther asunder, than they are in the *Newtonian Reflector*; for there is no Diffipation of Rays in the Reflection of Light from a Speculum, whose two Surfaces are parallel, (364.) and consequently excepting the Vicinity of the two Images, there can be nothing to hinder the Perfection of Vision by the *Newtonian Telescope*, and that is nothing compared to the Trouble of viewing Objects thro' the Side, which is the chief Reason of its being so little in Use.

377. Besides, were it worth while, we have a Theorem in Dr. *Gregory's Dioptrics*, by which the two Surfaces AB, CD, may be such as shall form both the Images in the same Part of the Axis; but then there will be the common Error from the Diffipation of Rays, and we lose more than we get by this Artifice.

378. Upon the whole it appears, that the REFLECTING TELESCOPE, truly made, is by far the most perfect of all dioptric Instruments; and is in its own Nature capable of every Improvement from Art, and Conveniency of Application, for viewing distant Objects in the most satisfactory and advantageous Manner.

379. As the double Refraction in the Speculum is made at a concave Surface of Glass AB, the *Error arising from the Figure* is here also capable of Correction so far as to render it insensible ; so that by this Contrivance both the Errors are provided for, (one wholly removed, the other sufficiently corrected) by reason the refracting Surfaces of the Speculum and Lens are of the *same Kind* ; whereas in two Lenses of a different Figure, tho' one Error be taken away, the other is greatly encreased ; so that we may esteem this Method of the *Speculum and Lens* as the *most perfect that can be in Point of Theory*, for removing and correcting of Errors in *dioptric Vision*.





C H A P. XIII.

LENSES made according to the Figures of the CONIC SECTIONS, or any other Curves, are subject to the same ERRORS from REFRACTION as those of a Spherical Form.

380. **A**S we have given a sufficient Demonstration that no *refracting Telescope* will admit of perfect and distinct Vision by *spherical Lenses*; let us next enquire what may be expected from other Forms; and here the CONIC SECTIONS next offer themselves.

381. The Theorem $\frac{\frac{1}{2}pnd}{md - nd - \frac{1}{2}pn} = f$, is general for finding the focal Distance of a Lens made with the Curvature of the *Ellipses, Parabola, or Hyperbola*, supposing $p =$ Parameter of each respectively, as I have elsewhere shewn*. But here it must be understood to relate to those Rays only, which fall extremely near the Axis.

Q 2

382.

* *Philosophia Britannica.*

The ABERRATION of CONIC LENSES

382. For none of those Rays which diverge from a Point at a finite Distance d , and are incident at any sensible Distance from the Axis, can by those figured Lenses be united again in a Point; nothing but a *mechanical Sort of Oval* being sufficient for this Purpose, of which *Descartes* has described four different Forms in his *Geometry*; and *Sir Isaac Newton* (by a general Construction) in *Prob. XXXIV.* of his optical Lectures. But those Forms are so irregular, that nothing can possibly be done with them in Practice.

383. There is indeed one Case, *viz.* where d is *infinite*, or the Rays *parallel*, that will admit of their being united into one Point in the Axis; and then the Theorem becomes $\frac{\frac{1}{2} p m}{m-n} = f$; where we observe the Theorem is the same as in a SPHERICAL Lens, whose Radius $r = \frac{1}{2} p$, the Semi-parameter of the conic Section (340.)

384. But these parallel Rays are to be understood to be also *parallel to Axis* of the Section or Lens that is made thereof; for those that flow from a Point at an infinite Distance, but out of the Axis, will not be united into a Point by such a *conic Lens*; and therefore after *Descartes* had himself made many compound Lenses of these Forms, and applied them in every different Way, he at last concludes, —“ That the Inequality of Curvature in these Lines (or Sections)
“ which

“ which compose the Figure of these Lenses,
 “ will ever be the Cause why Rays which are
 “ referred to divers Points, or parallel Rays
 “ proceeding from divers Parts, can never be
 “ so equally and exactly refracted as those are
 “ which respect one Point only, or come from
 “ one and the same Part parallel upon the
 “ Lens *.”—Dr. BARROW, who published his
 optical Lectures soon after, declared he had
 nothing to add on this Head after *Descartes*; and
 the same is intimated also by Sir *Isaac Newton*.

385. If therefore no equal and uniform Refraction of Rays, which come from divers Points, can be effected by *conic Lenses*, there must be an *Aberration* of those Rays, and such as cannot be corrected by any Sort of *Combination*. The Image therefore of an Object will necessarily be imperfect in the Focus of such a Lens, single or compound, and consequently no refracting Telescope can admit of perfect or distinct Vision, if constructed with such Lenses.

386. But supposing the Figure of those Lenses would form a perfect Image, yet that would stand us in little stead, while the Means of attaining to those Figures are insuperably difficult, if not altogether impossible; for supposing the focal Distance of the convex Object-Glass were only 9 Inches, then in a Lens of the *elliptic, parabolic, hyperbolic, and spherical* Form, the Radius

* Dioptrics, Page 122.

The ABERRATION *from the* FIGURE

dius of Curvature at the Vertex of each is the same, *viz.* 3 Inches, equal to the Semi-parameter of those Curves; and at that small Distance on each Side the Axis, as is equal to the Semi-diameter of the Object-Glass, (*viz.* $\frac{1}{2}$ an Inch) all those Curves would so nearly coincide, as by no Means to admit of any sensible Difference; and consequently leave us without any Possibility of making Tools of such an exact and determinate Figure as is necessary to produce the Lenses under Consideration. In this Sort of *Enchiridion* a Man may possibly equal, but can never hope to exceed *Descartes*, who after all his tedious and laborious Endeavours, was obliged to desist, finding the Laws of Refraction as irresistible as those of Fate itself.

Since then it appears that Lenses formed according to the *conic Sections*, or any other Curve whatsoever, are subject to the same Kind of Refraction from the Figure as spherical Lenses are, it is evident, it must be corrected after the same Manner likewise, *viz.* by making the optic Angle by two Refractions instead of one; and consequently both the Refractions must be the same Way, or made with two convex or two concave Lenses; for if the Refractions are contrary, or made by a *convex or concave Lens* conjointly, instead of correcting they must encrease the Error greatly, as was shewn to be the Case

Of CONIC LENSSES *not to be annihilated.*
Case of those of a spherical Figure. If any
Doctrinē be advanced to the contrary of this, the
*Public has a Right to demand a Demonstration of
the Truth of it;* till then, they may think the
achromatic Telescope itself a sufficient Proof
that no *refracting Telescope* can be so *distinct* as a
Reflector. They will also most of them be so
prudent, as to remember the Motto, *Nullius
in Verba,* &c.

119

F I N I S.



Addenda, Corrigenda, &c.

It may be proper to remark, that whereas there has been of late Years great Talk of a *Night Telescope*, as if it were a modern Invention, the Reader is to know that this Telescope is in reality no other than the *astronomical Telescope with three Glasses*, the Invention of which Construction was before the Time of Mr. *William Molyneux*, who in his *Dioptrics*, speaking of a Combination of two Lenses, has this Remark, —*This Problem (viz. of finding the Focus of the Glasses so combined) is of considerable Use in Dioptrics, being the Foundation of an excellent Sort of Telescope much used in England for the Night.* Page 73. This Book was dedicated in the Year 1690. So that this *Night Telescope* has been in Use at least 70 or 80 Years. See also in *Huygens's Dioptrics* the Construction of the same Telescope. Page 184.

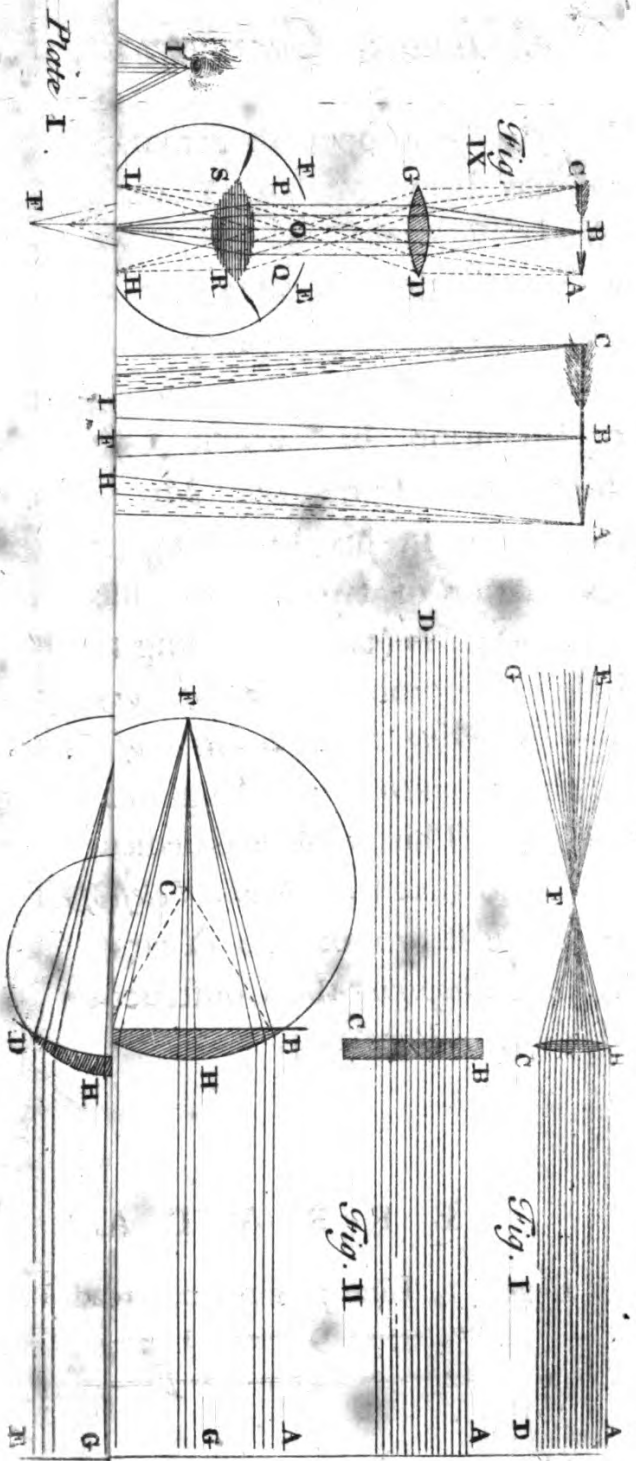


E R R A T A.

Page 41, Line 7, for $x=2$ read $x=1,2$:

ib. — 9 — 1 — 1,8.

42 — 24 — ~~ff~~ — +ff.



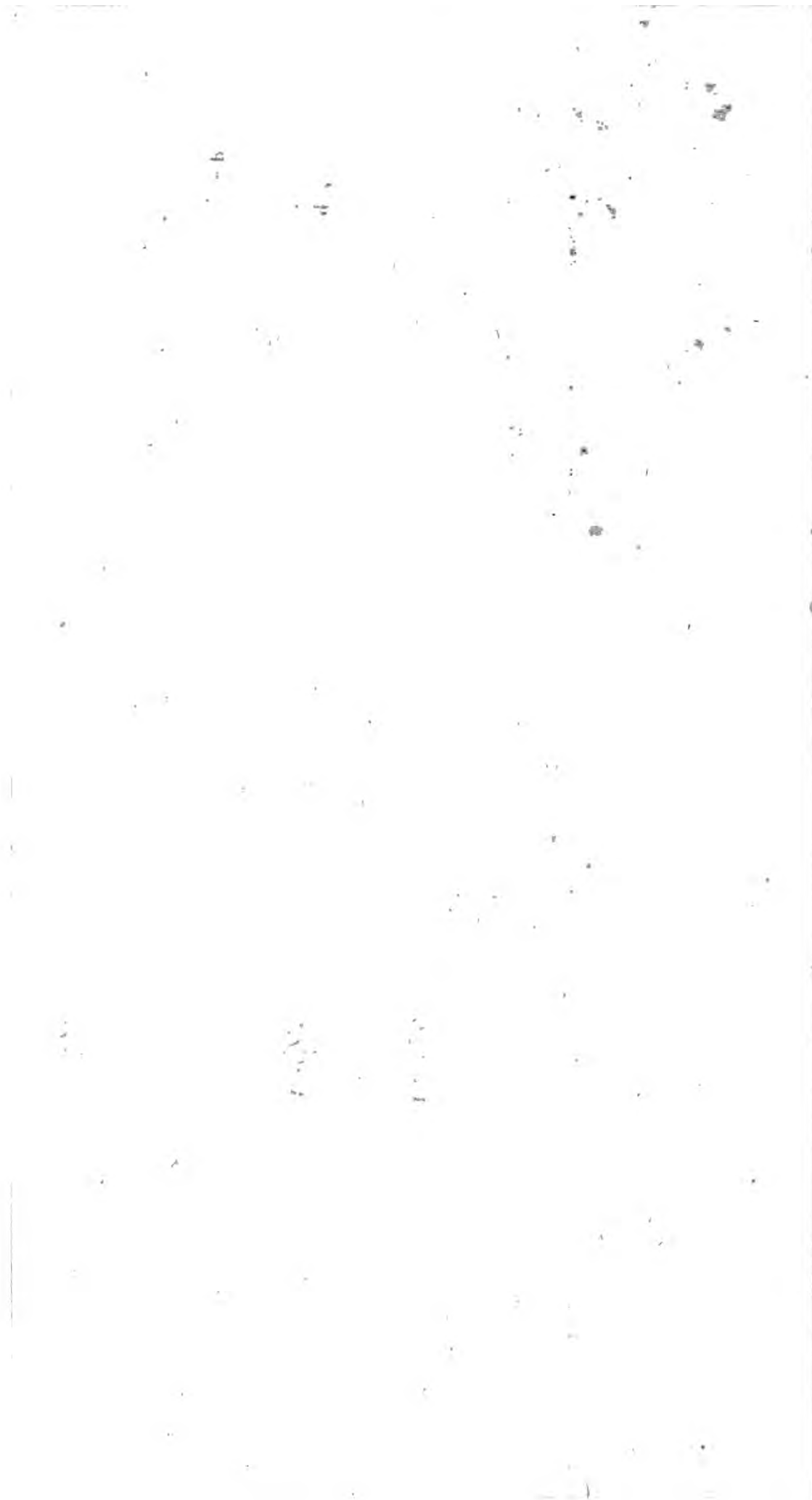
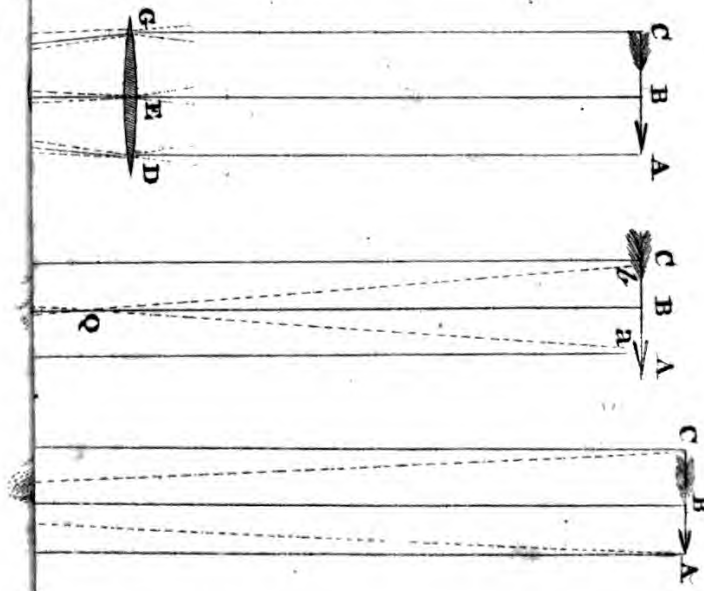
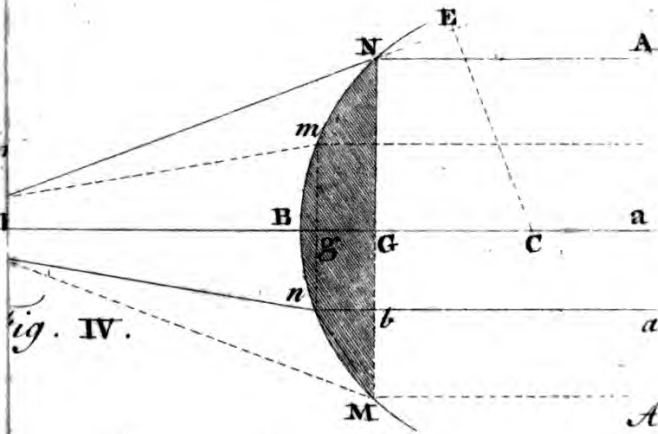
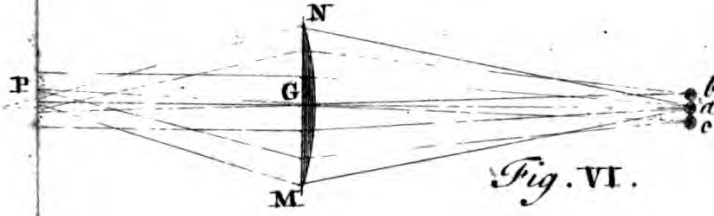


Plate II.





M I P M I R

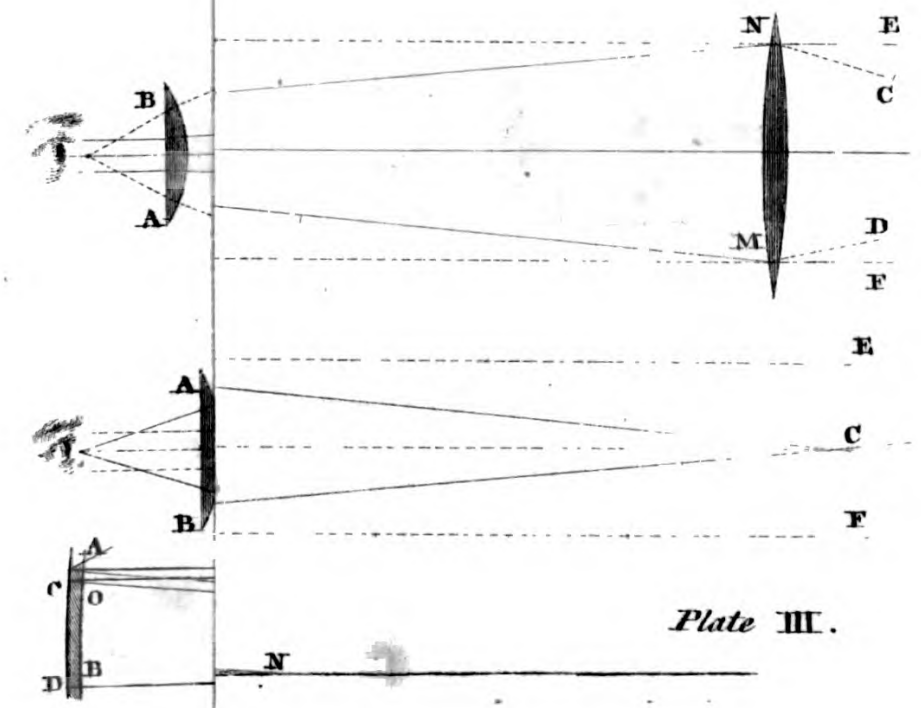
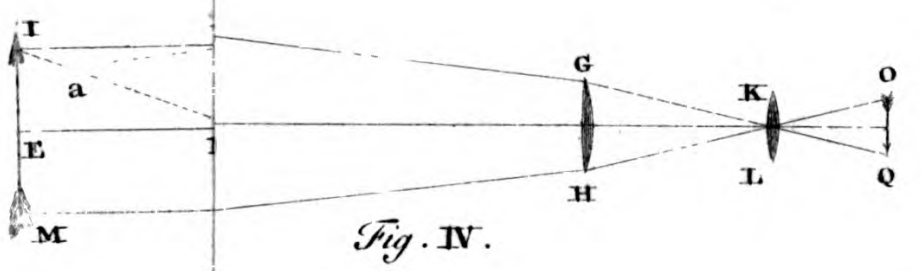
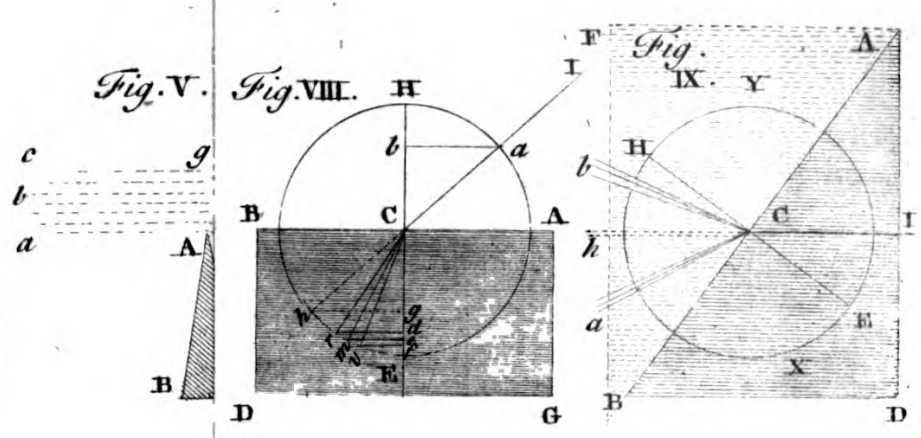


Plate III.

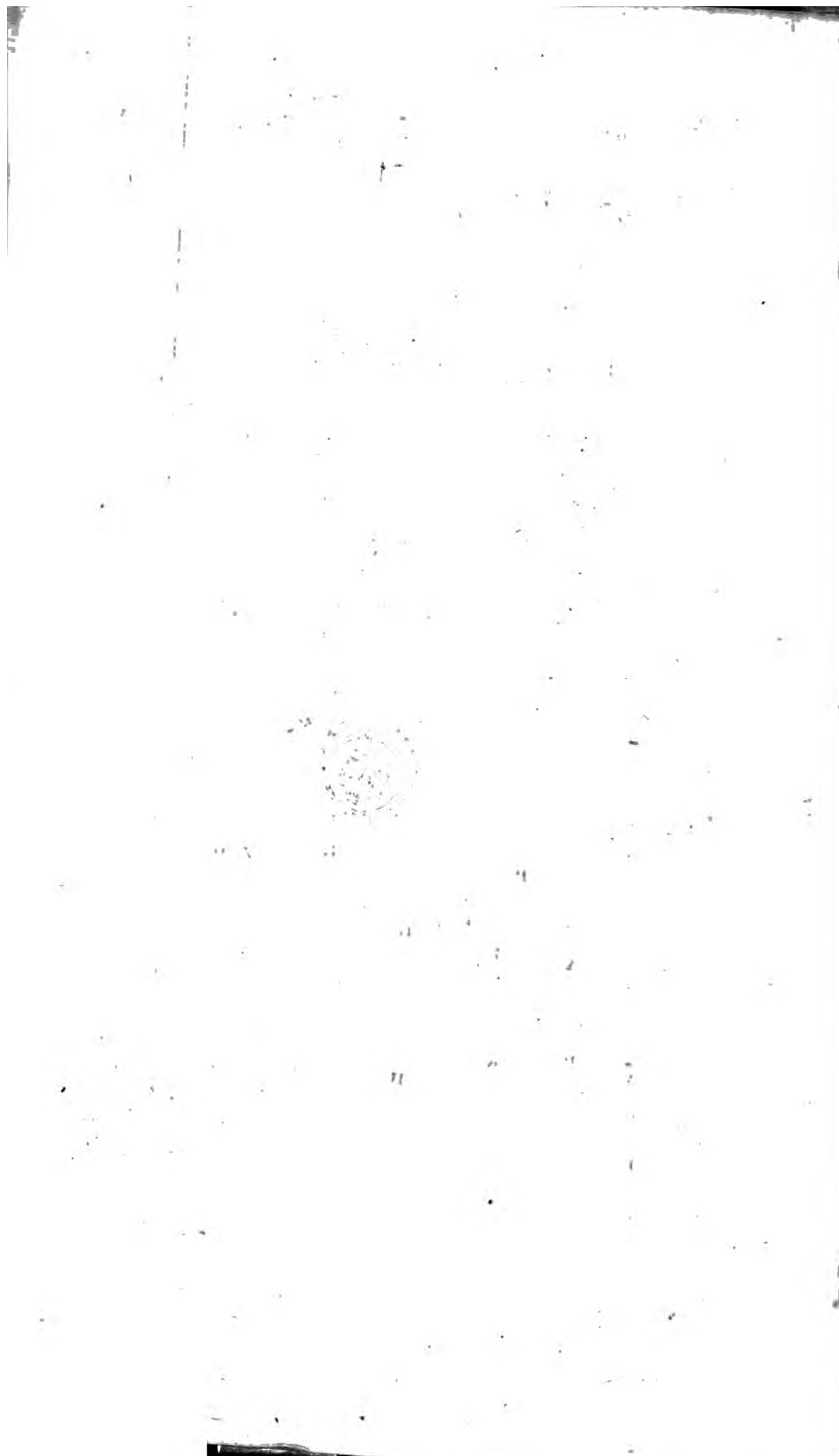


Plate IV. Fig. I.

Fig. I.

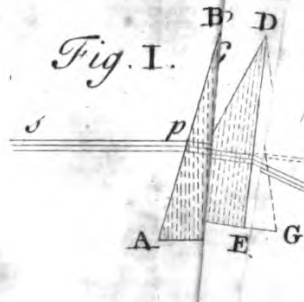


Fig. V.

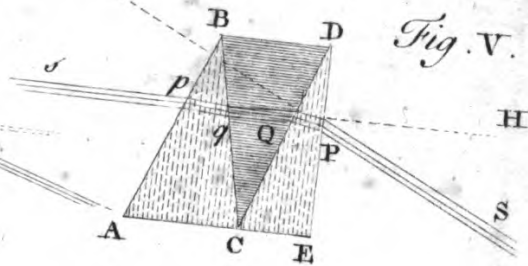
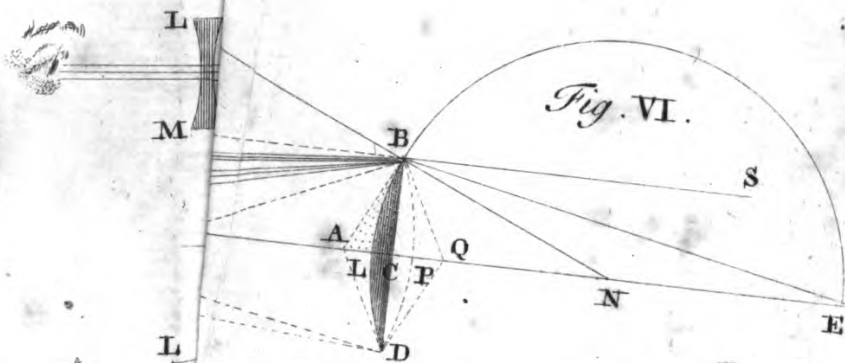


Fig. VI.



N VIII.



Fig. X.

